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Intro to SR & FEL spectroscopy; interaction between radiation & matter (2)

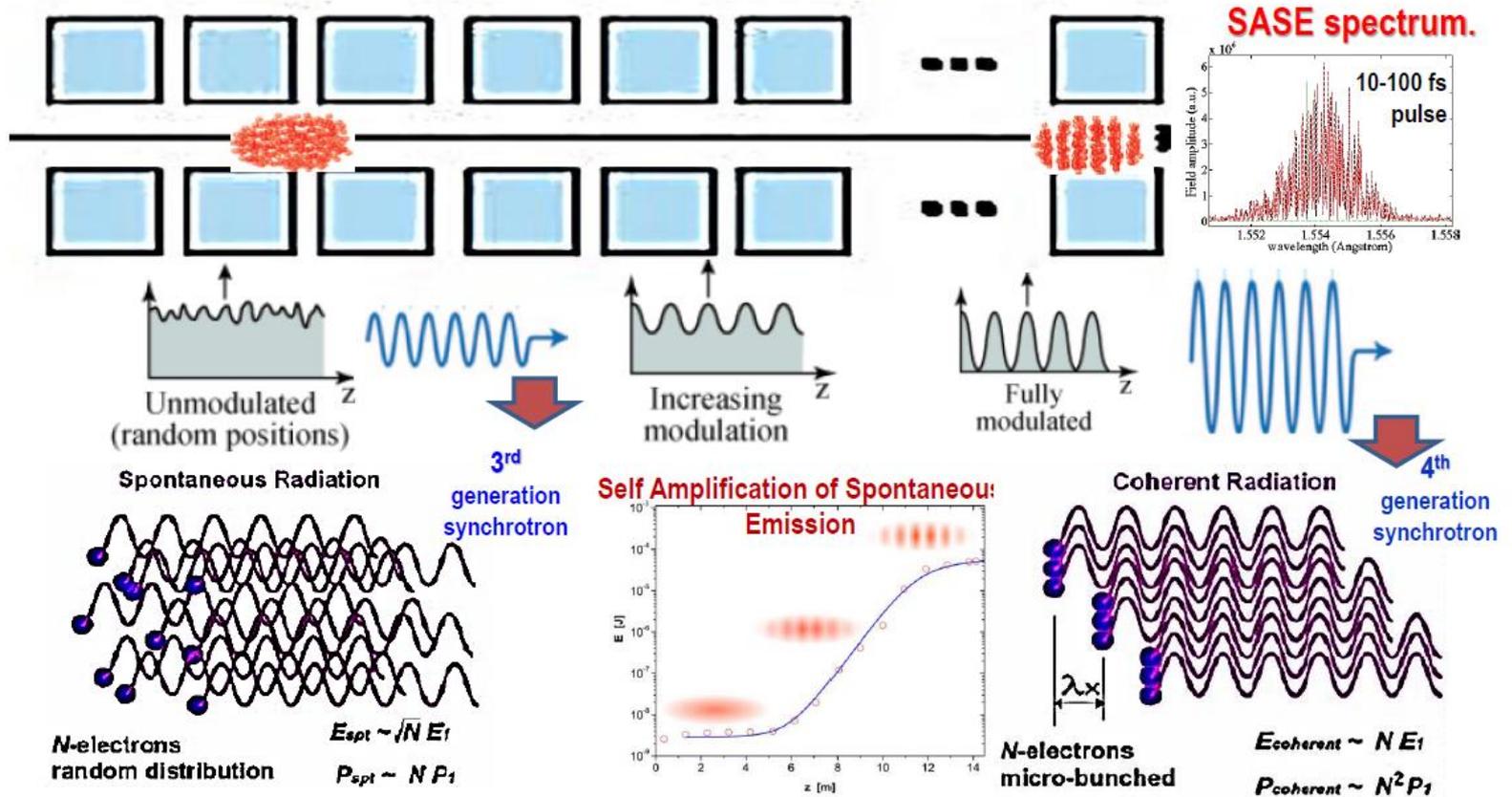
Maria Novella Piancastelli

*Sorbonne Universités, UPMC Univ Paris 06, CNRS, Laboratoire de Chimie
Physique-Matière et Rayonnement, France*

*Department of Physics and Astronomy,
Uppsala University, SE-75120 Uppsala, Sweden*



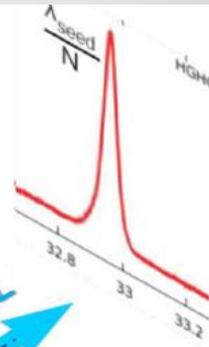
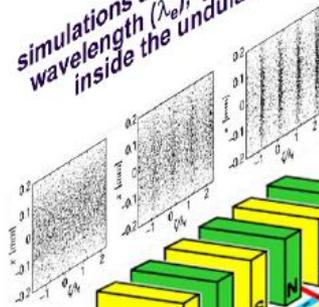
With a very long undulator the radiated fields become stronger and lead to microbunching, i.e. transform the random positions and motions of electrons into correlated waves of electrons, emitting radiation in phase.





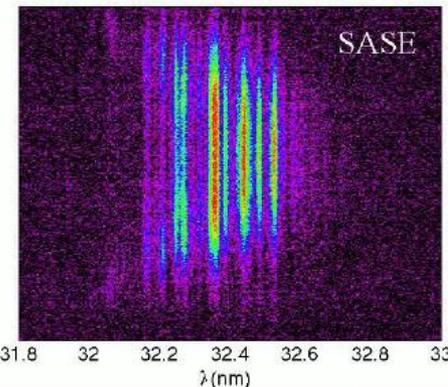
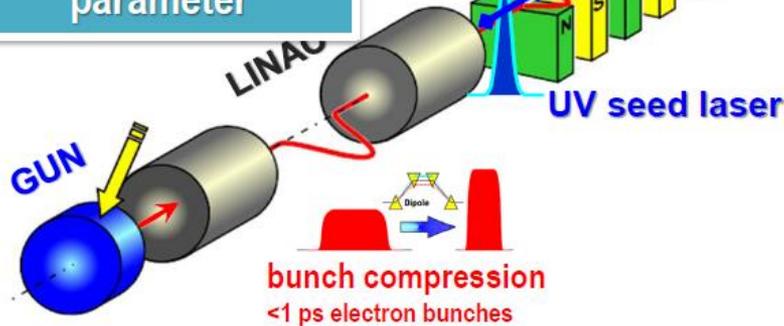
High Gain Harmonic Generation (HGHG): seeding (modifying) the emitting electron bunch with an external laser pulse controlled in all the relevant photon parameter

simulations at the radiation wavelength (λ_0), ξ - distance inside the undulator



“SASE” FEL – several separate “waves” of electrons with uncorrelated phase. Less peak power, broader spectrum.

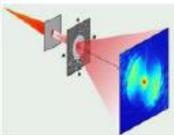
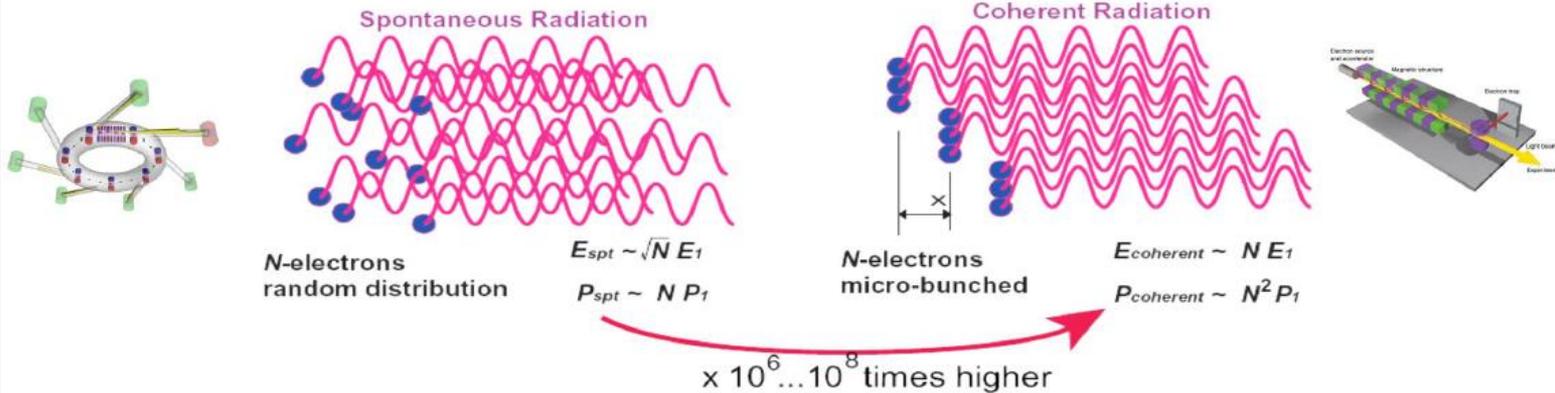
The properties of the FEL radiation are entangled with those of the seed laser. Defined energy-time profile.



Synchrotron vs FEL Radiation

long pulses (sub-ns) max $\sim 10^8$ photons/pulse

Fs pulses $> 10^{11}$ photons/pulse

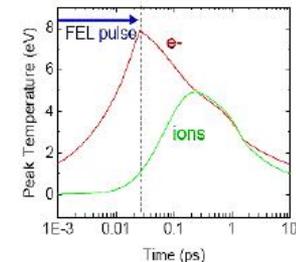


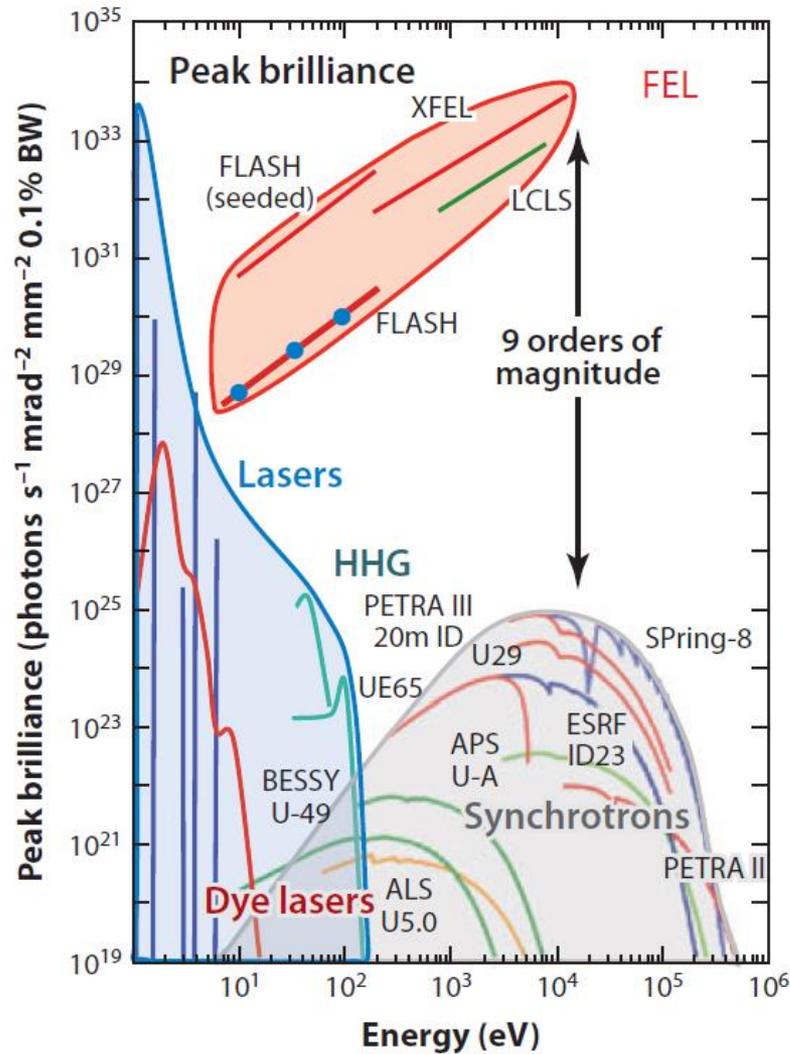
Synchrotron radiation:
pinhole and monochromators for spatial and spectral filtering, but at the expense of intensity!



FEL (FLASH, LSLs, SACLA, FERMI):
natural space coherence: each electron - spontaneous emission that overlap each other in phase

Ultra-short (fs) and ultra-bright coherent FEL pulses allow imaging with single pulse before the radiation damage manifests itself !







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	LCLS	LCLS II	Eu-XFEL	SACLA	FLASH	FLASH II	FERMI	SwissFEL	PAL XFEL	Shanghai XFEL
Shortest wavelength	1.5 Å	1 Å	0.5 Å	1 Å	40 Å	40 Å	40 Å	1 Å	1(0.6) Å	1 Å
Undulator type hard X-ray.	Fixed gap	Variable gap	Variable gap	In-vacuum Var. gap	n.a.	n.a.	n.a.	In-vacuum var. gap	Variable gap	Variable gap
Undulator type soft X-ray.	n.a.	Variable gap	Variable gap	n.a.	Fixed gap	Variable gap	Apple II	Apple II	Apple II	?
Injector	S-band RF gun	S-band RF gun	L-band RF gun	Pulsed Diode	L-band RF gun	L-band RF gun	S-band RF gun	S-band RF gun	S-band RF gun	S-band RF gun
Cathode	Cu	Cu	Cs ₂ Te	CeB ₆ (thermionic)	Cs ₂ Te	Cs ₂ Te	Cu	Cu	Cu	Cu
Main linac technology	n.c. Pulsed	n.c. pulsed	s.c. pulsed	n.c. pulsed	s.c. pulsed	s.c. pulsed	n.c. pulsed	n.c. pulsed	n.c. pulsed	n.c. pulsed
RF frequency	S-band	S-band	L-band	C-band	L-band	L-band	S-band	C-band	S-band	C-band
RF Rep. rate	120 Hz	120 Hz	10 Hz	60 Hz	10 Hz	10 Hz	10-50 Hz	100 Hz	120 Hz	60 Hz
FEL pulses/RF pulse	1	1	2700	1	2700	2700	1	2	1	1
max. bunch charge	0.25 nC	0.25 nC	1 nC	0.2 nC	1 nC	1 nC	0.5 nC	0.2 nC	0.2 nC	0.2 nC
max. electron energy	13.6 GeV	14 GeV	17.5 GeV	8 GeV	1.2 GeV	1.2 GeV	1.5 GeV	5.8 GeV	10 GeV	6.4 GeV
No. RF stations	81	81	29	69	5	5	15	34	49	?
Approx. facility length	1.7 km	1.7 km	3.4 km	0.8 km	0.32 km	0.32 km	0.5 km	0.7 km	1.1 km	0.6 km
Start operation	2009	2017	2015	2011	2005	2013	2010	2016	2015	2019



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The Physics of Free Electron Lasers

By Evgeny Saldin, Evgeny A. Schneidmiller, M.V. Yurkov
Springer, 2000

Classical Theory of Free-Electron Lasers

By Eric B. Szarmes
Morgan & Claypool Publishers, 2014

Insertion Devices for Synchrotron Radiation and Free Electron Laser

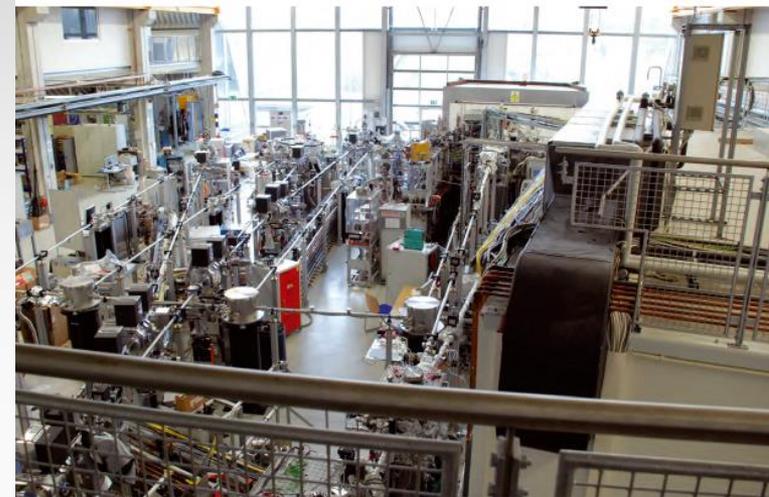
By F. Ciocci, G. Dattoli, A. Torre, A. Renieri
World Scientific, 2000



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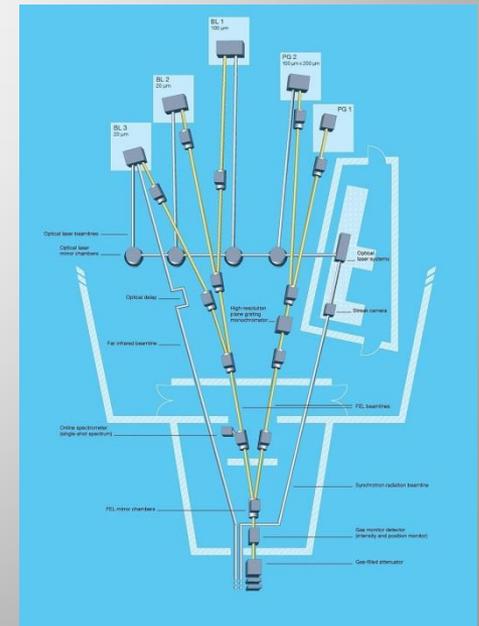
FLASH

DESY, Hamburg, Germany



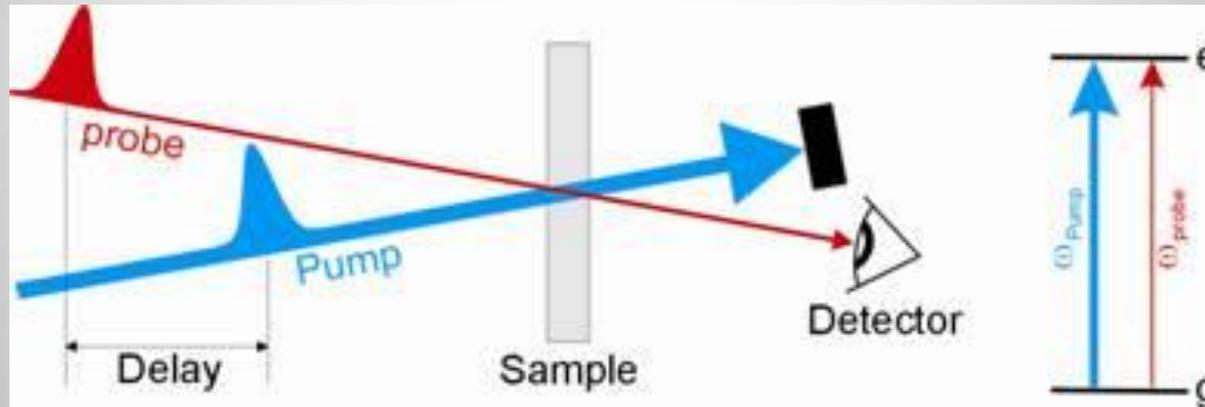
FLASH Parameters

XUV - Photons		
Number of Bunches per second+	1 - 8000 (delivered up to 5000 to users)	
Pulse Repetition Rates (within pulse train)+	40, 50, 100, 200, 250, 500, 1000	kHz
Wavelength	4.2 - 45	nm
Photon Energy	28 - 295	eV
Pulse Duration (FWHM)	30 - 300	fs
Average Pulse Energy (single bunch)*	1 - 500	μ J
Average Pulse Energy (pulse trains)*	1 - 200	μ J





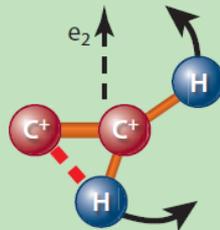
Pump-probe experiments



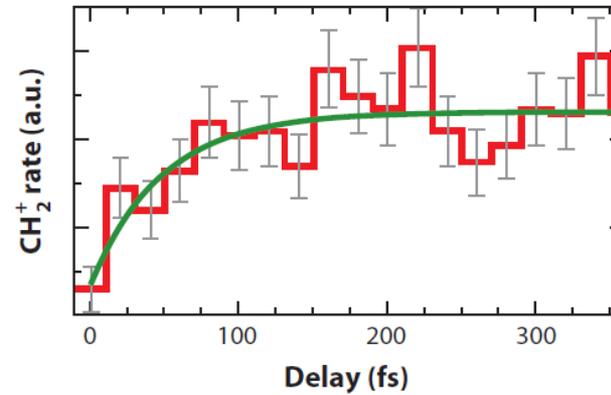
a Isomerization

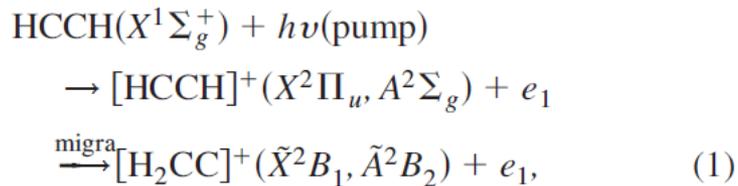
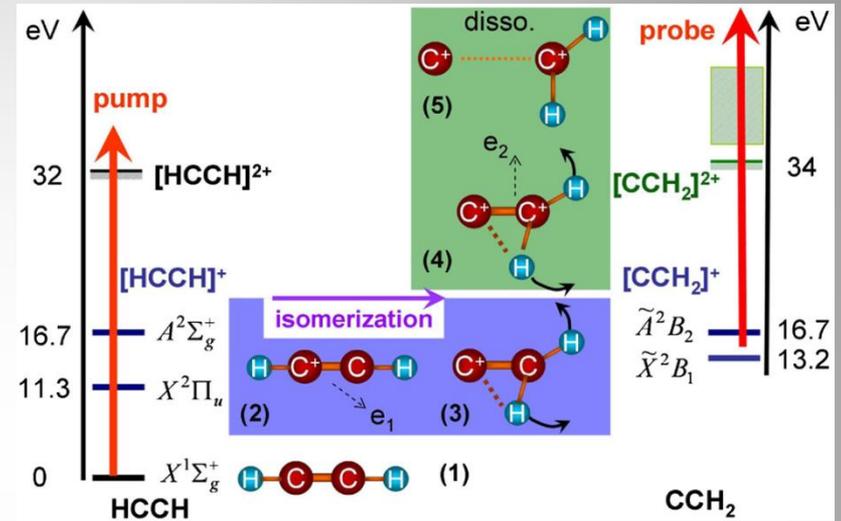


b

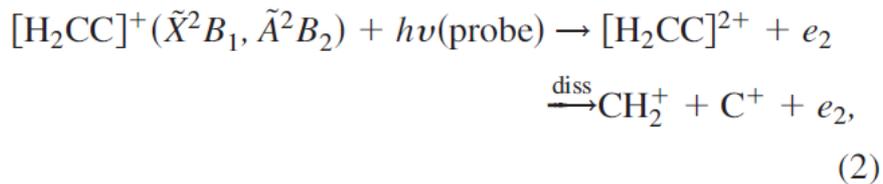


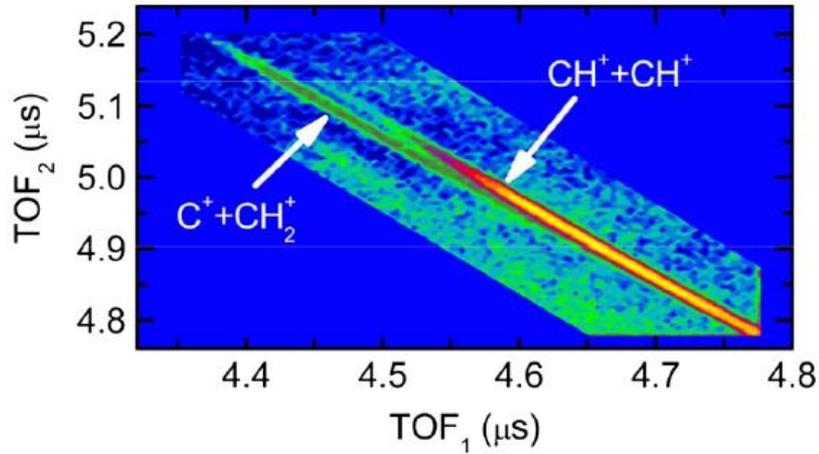
c





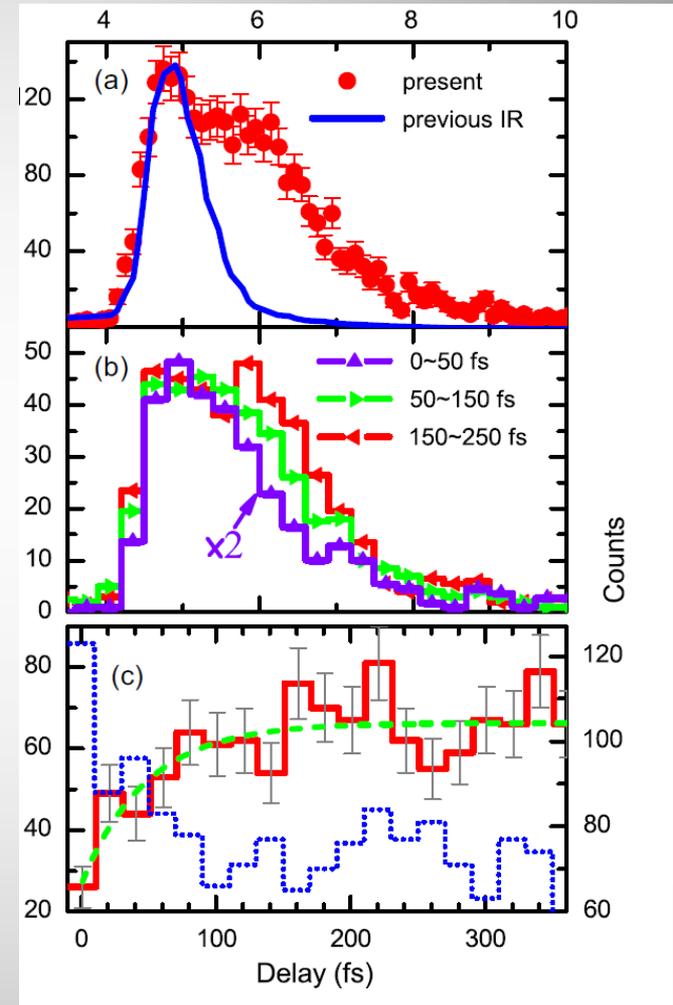
and





Mean isomerization time:

52 ± 15 fs





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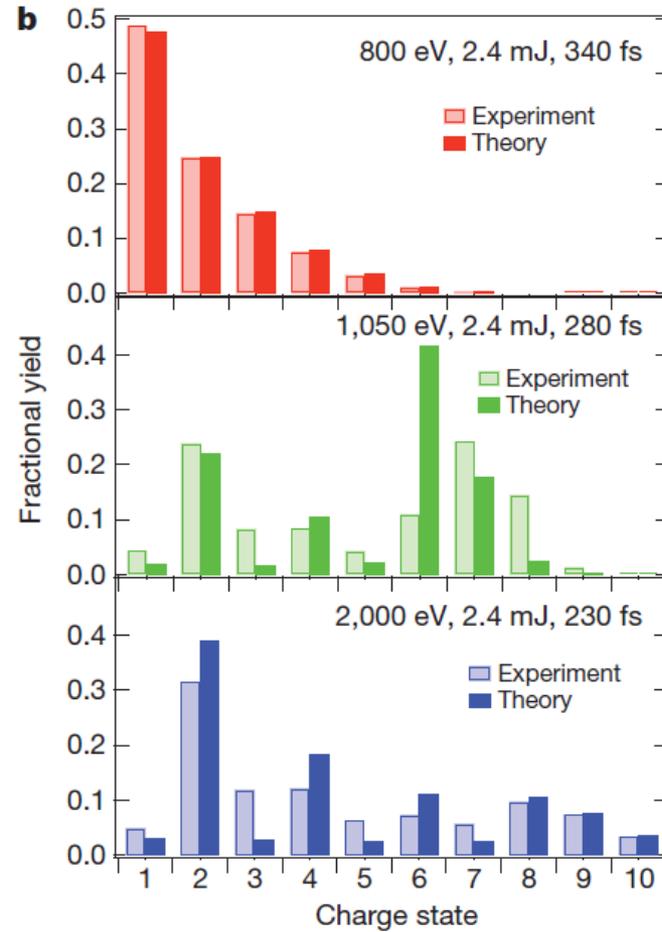
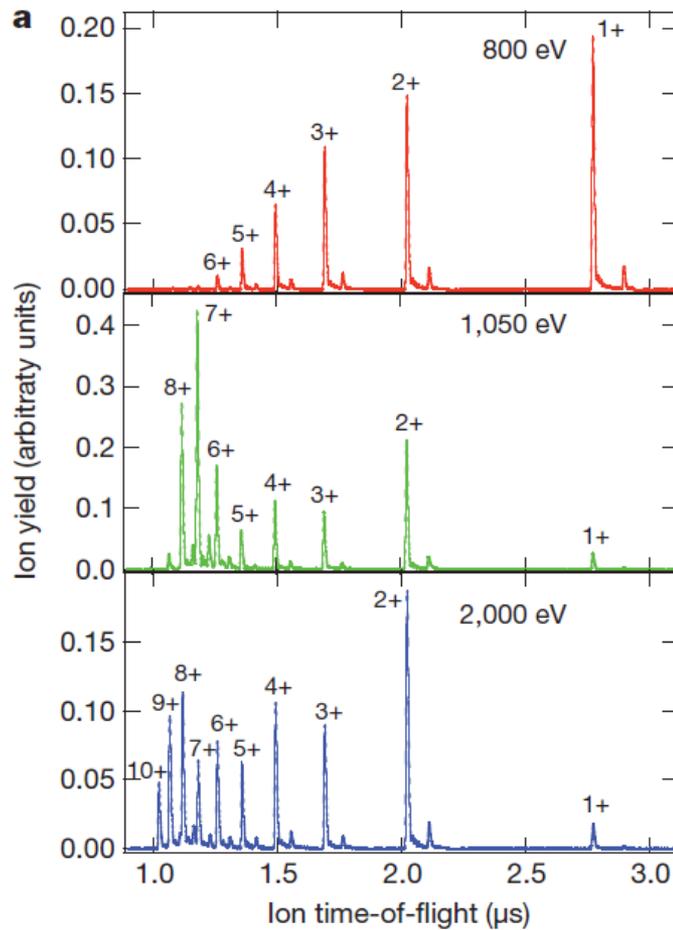


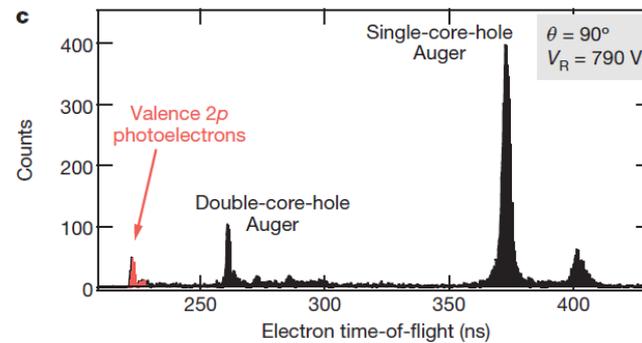
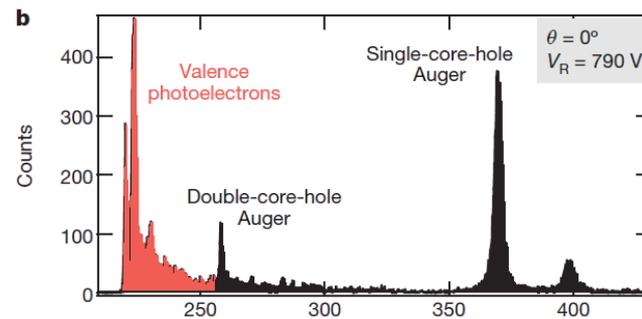
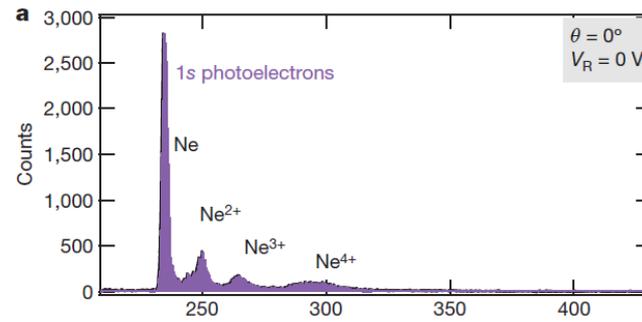
Photon Beam Parameters	Symbol	hard x-rays	soft x-rays	short pulse soft	short pulse hard	unit
Fundamental wavelength	λ_r	6.2-1.3	43.7-6.2	43.7-6.2	6.2-1.3	Å
Photon energy	$\hbar\omega$	2000-9600	285-2000	285-2000	2000-9600	eV
Final linac e^- energy	γmc^2	6.7-14.7	2.5-6.7	2.5-6.7	6.7-14.7	GeV
FEL 3-D gain length	L_G	3.3	1.5	~1.5	~3.3	m
Photons per pulse	N_γ	2	20	0.5	0.2	10^{12}
Peak brightness	B_{pk}	20	0.3	?	?	10^{32} §
Average brightness (120 Hz)	$\langle B \rangle$	160	8	?	?	10^{20} §
SASE bandwidth (fwhm)	$\Delta\omega/\omega$	~0.2-0.5	~0.2-1.0	?	?	%
Final pulse duration (fwhm)	$\Delta\tau_f$	50-250	70-400	<10	<10	fs

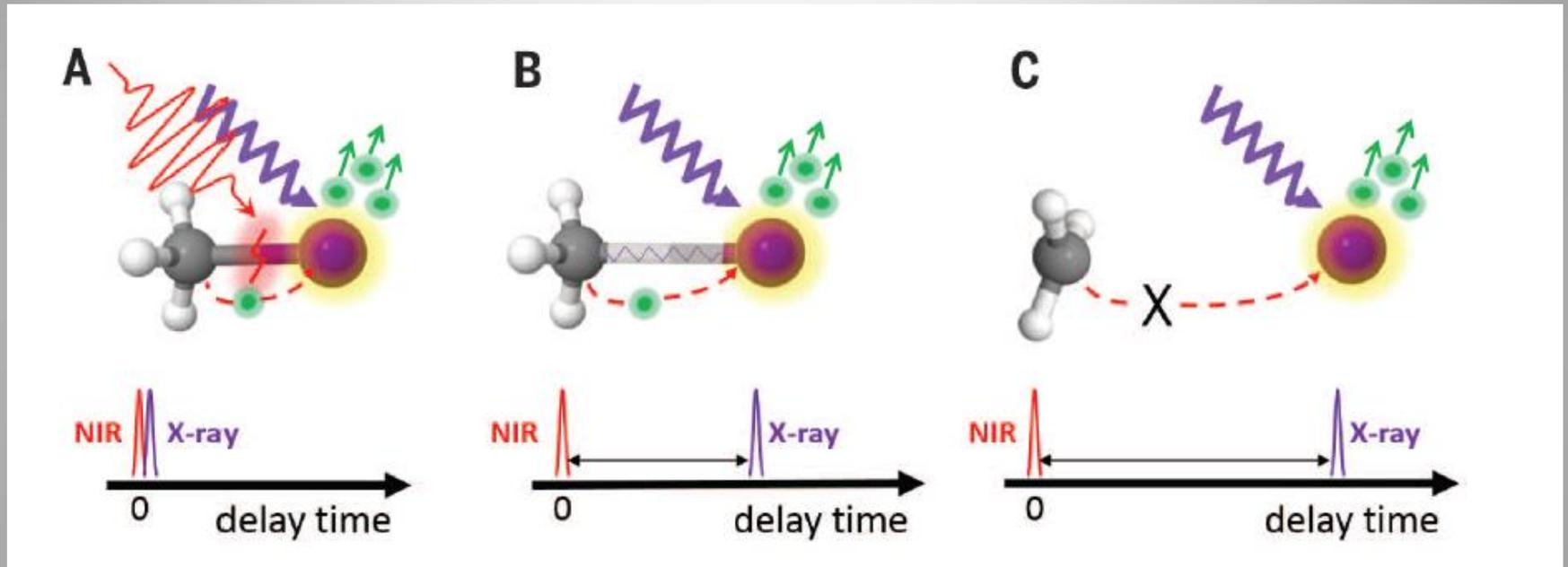
LCLS

SLAC, Stanford, CA, USA

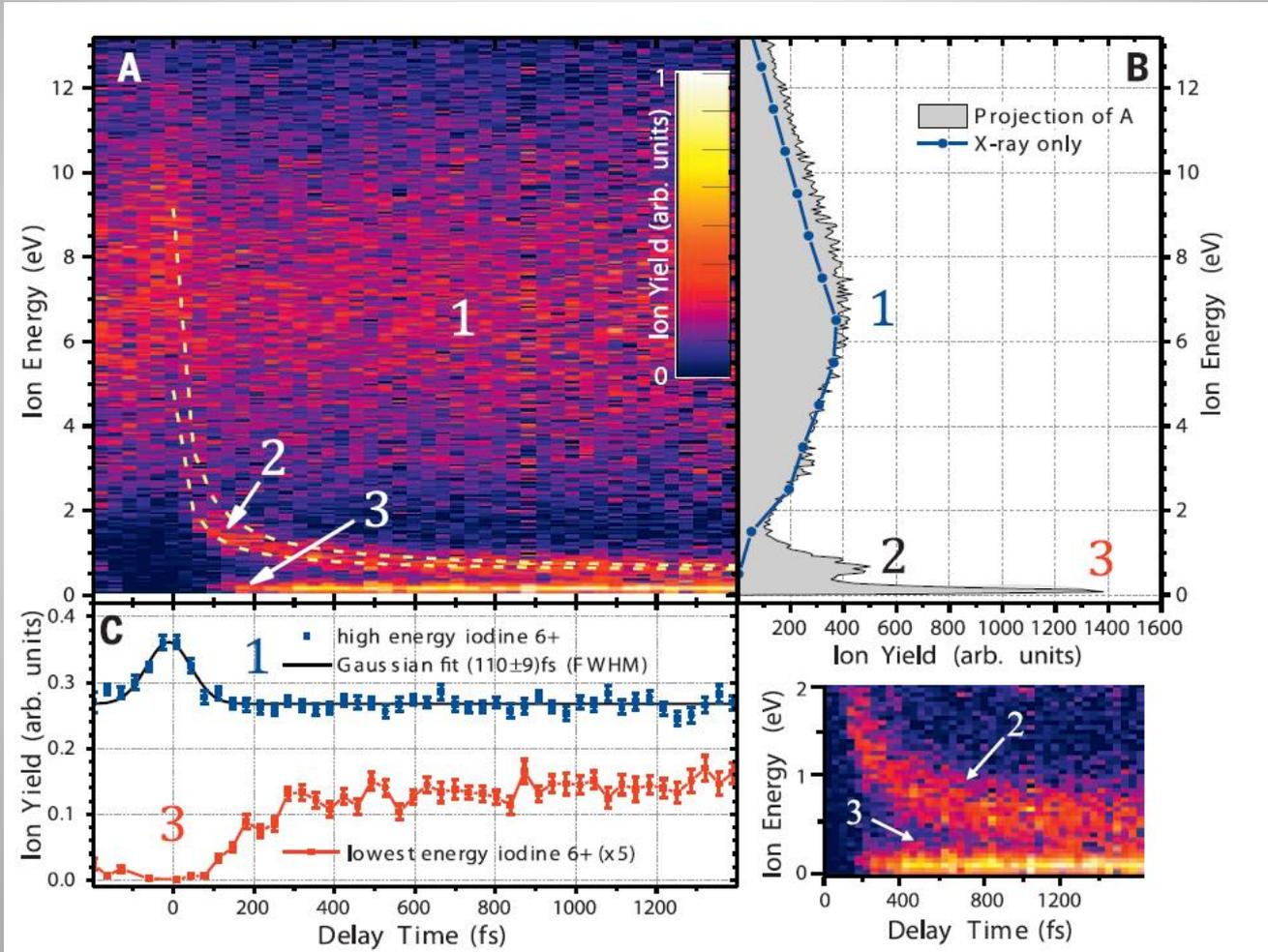


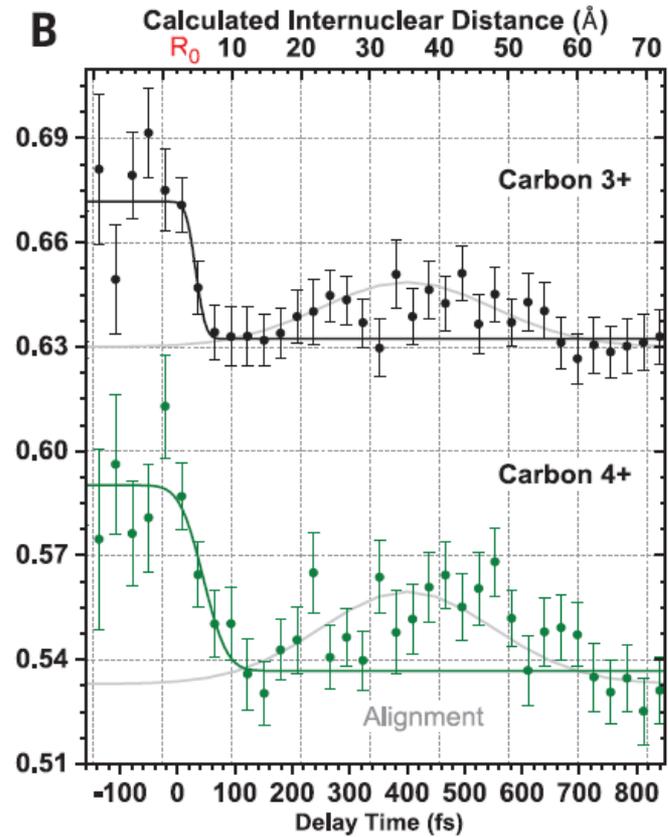
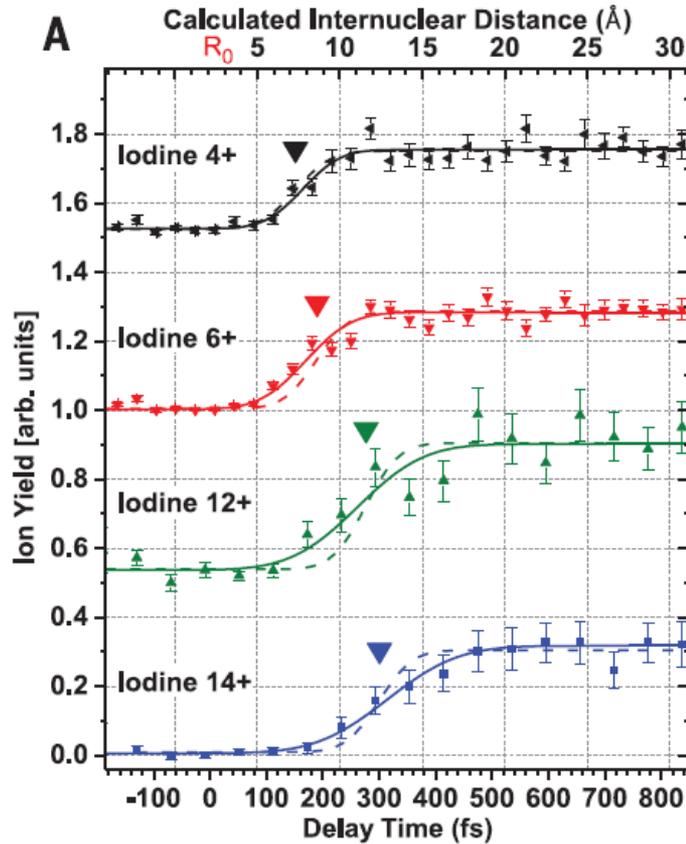






B.Erk et al., Science 345, 288 (2014)



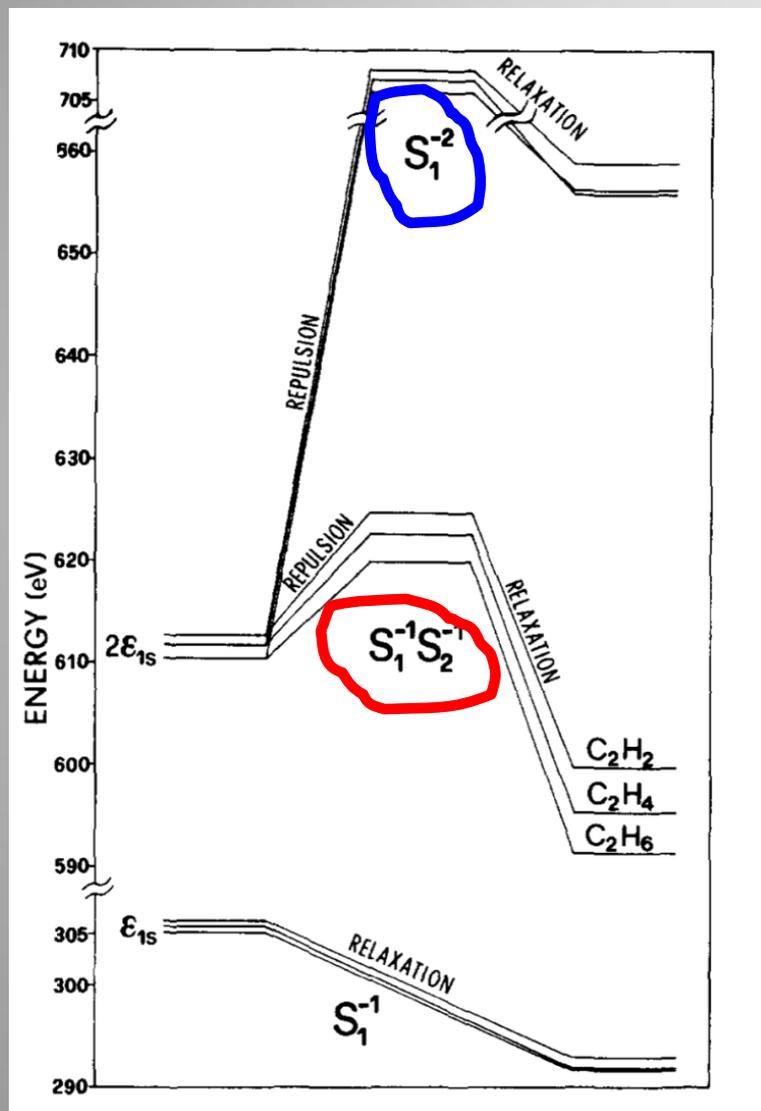




DCH Single-Site (SS)

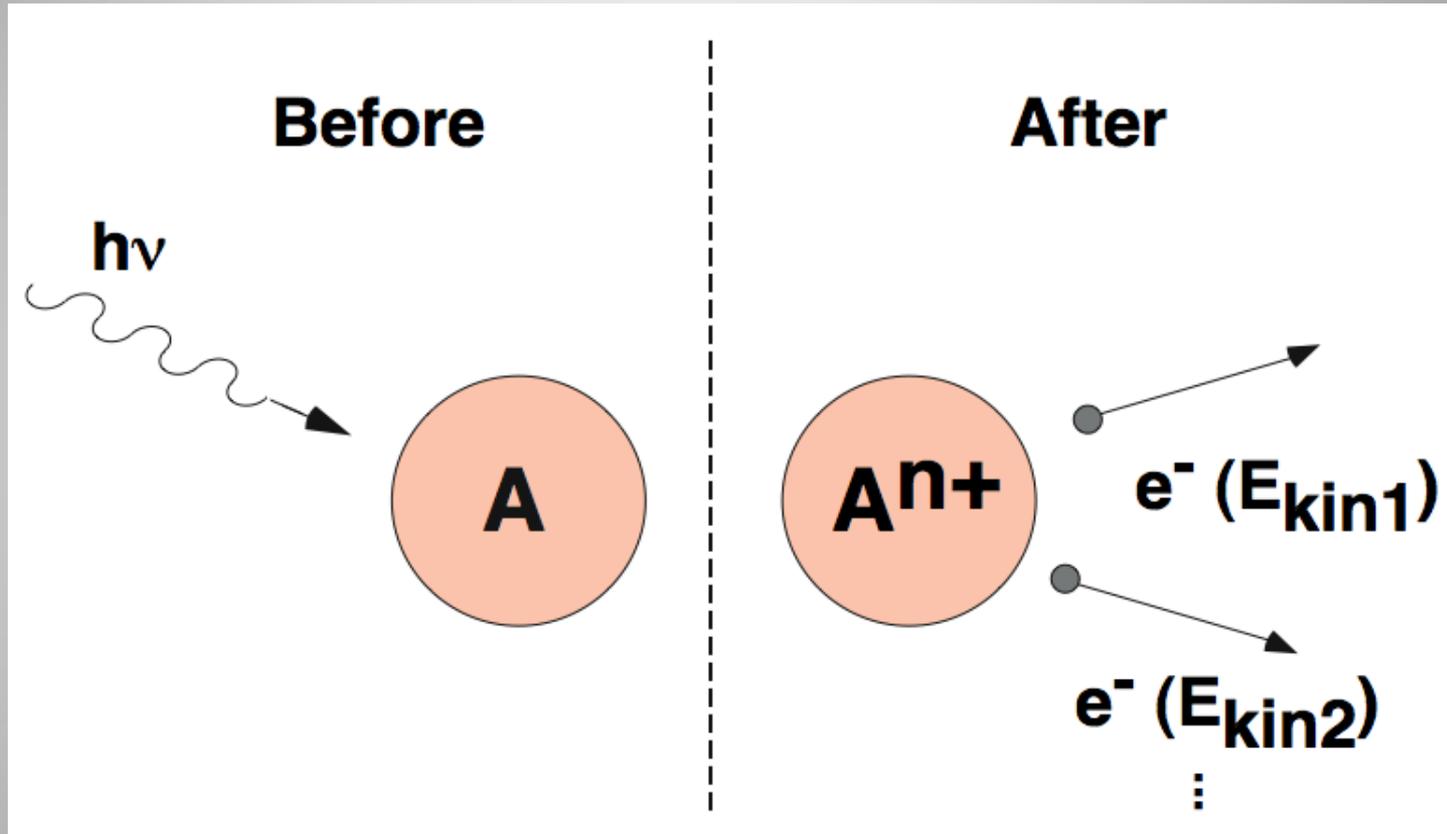
DCH Two-Site (TS)

Substantially enhanced
chemical sensitivity



L.S.Cederbaum et al, J.Chem.Phys. 85,
(1986) 6513

Single Photon – Multiple Ionisation

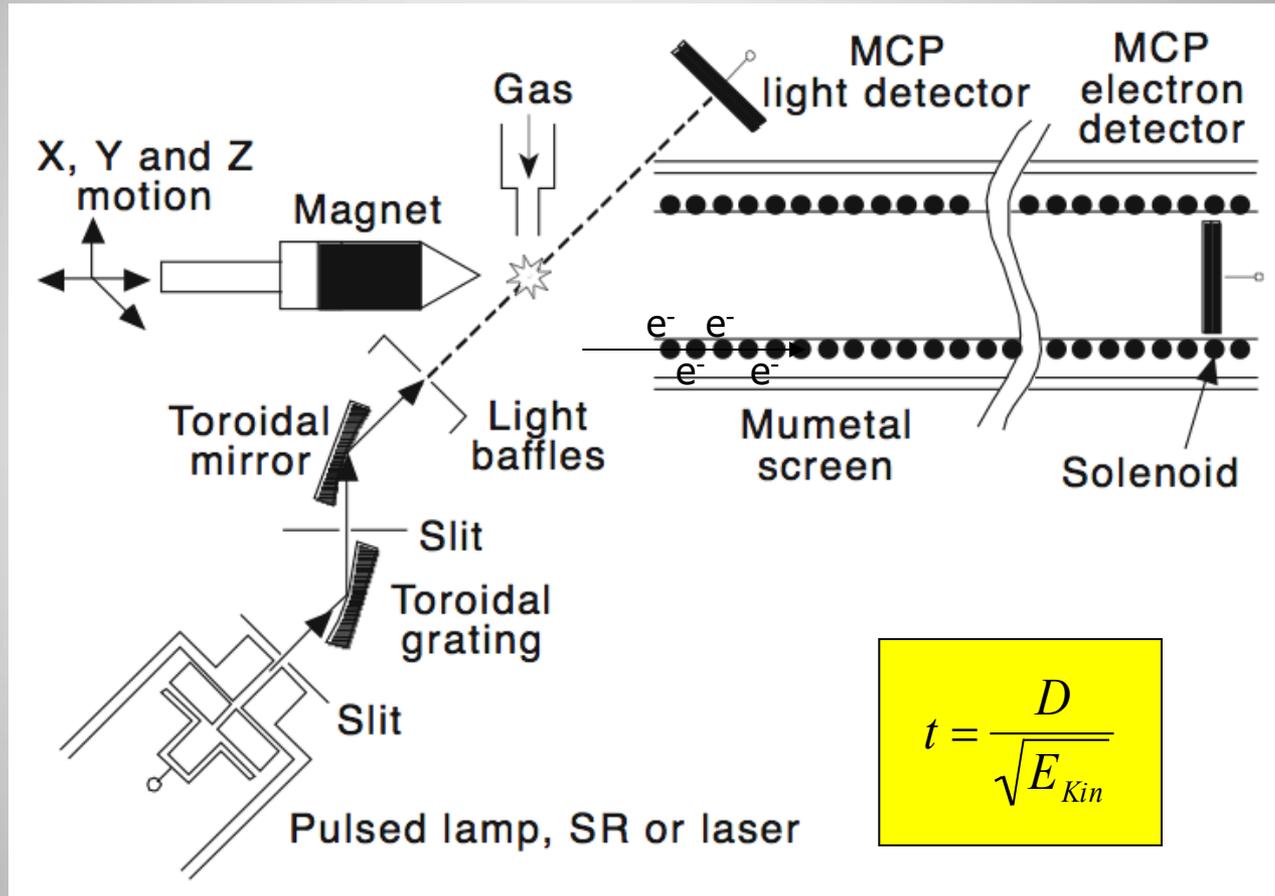


$$E_{B1} + E_{B2} (+ \dots) = h\nu - E_{kin1} - E_{kin2} (- \dots)$$

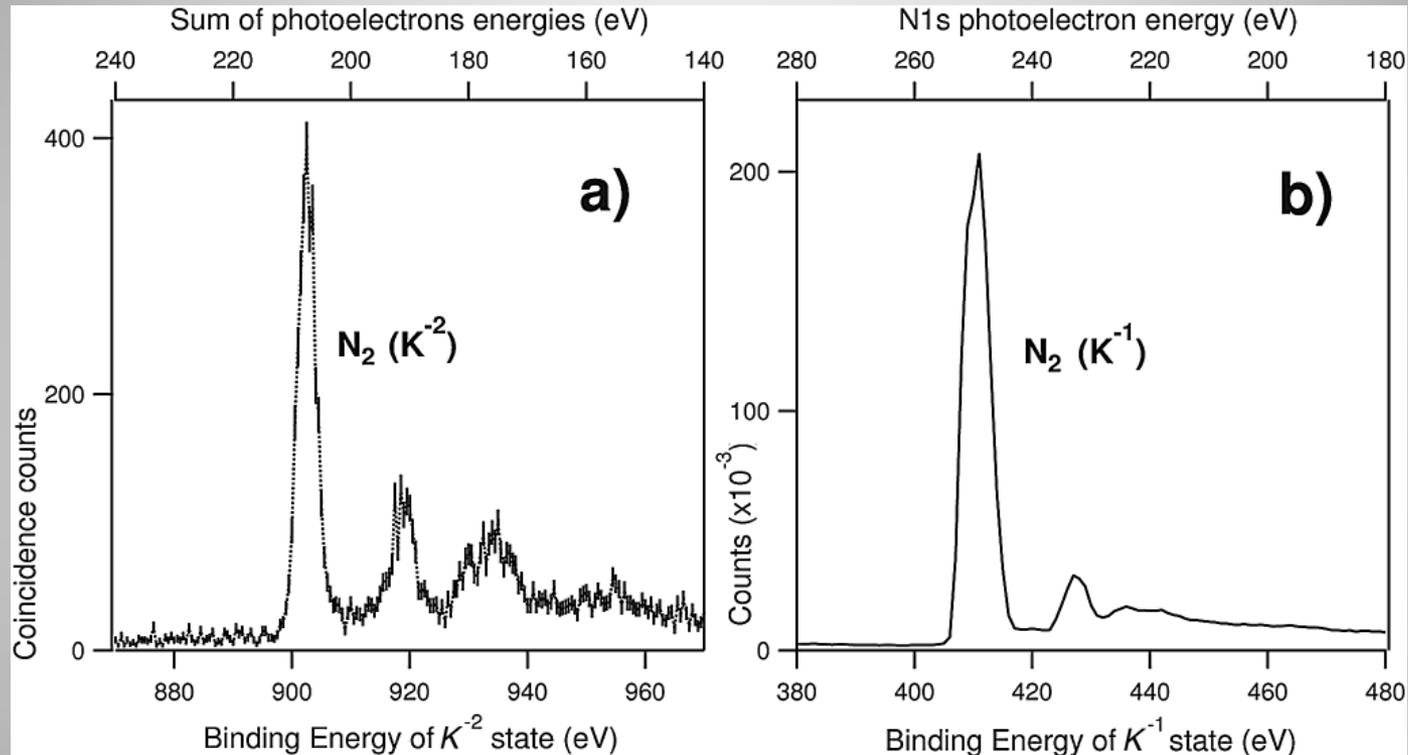
Double (multiple) ionisation energy (DIP, ...)

Time-Of-Flight Magnetic Bottle Multi-Electron Spectrometer

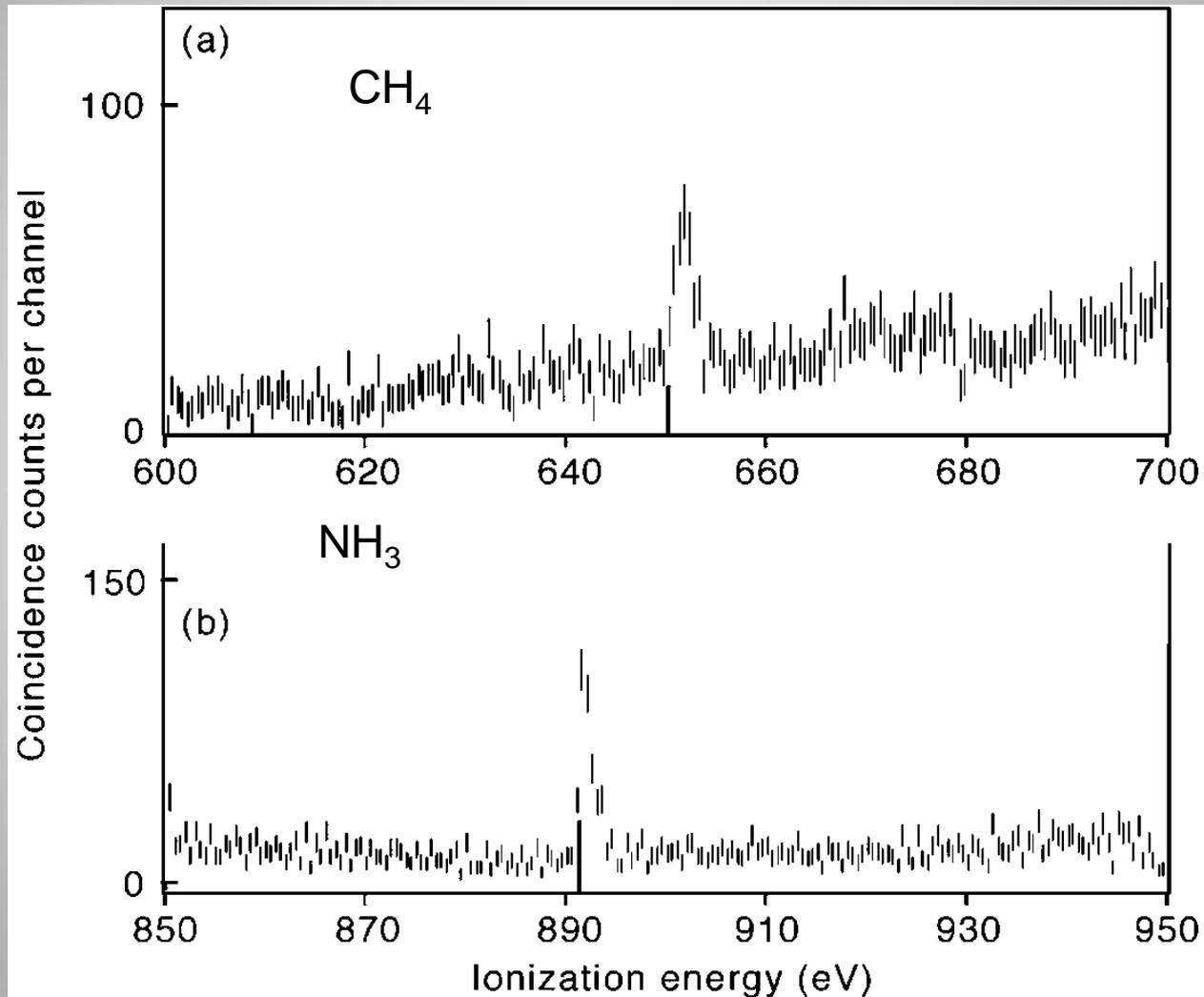
“4 π ” correlation device

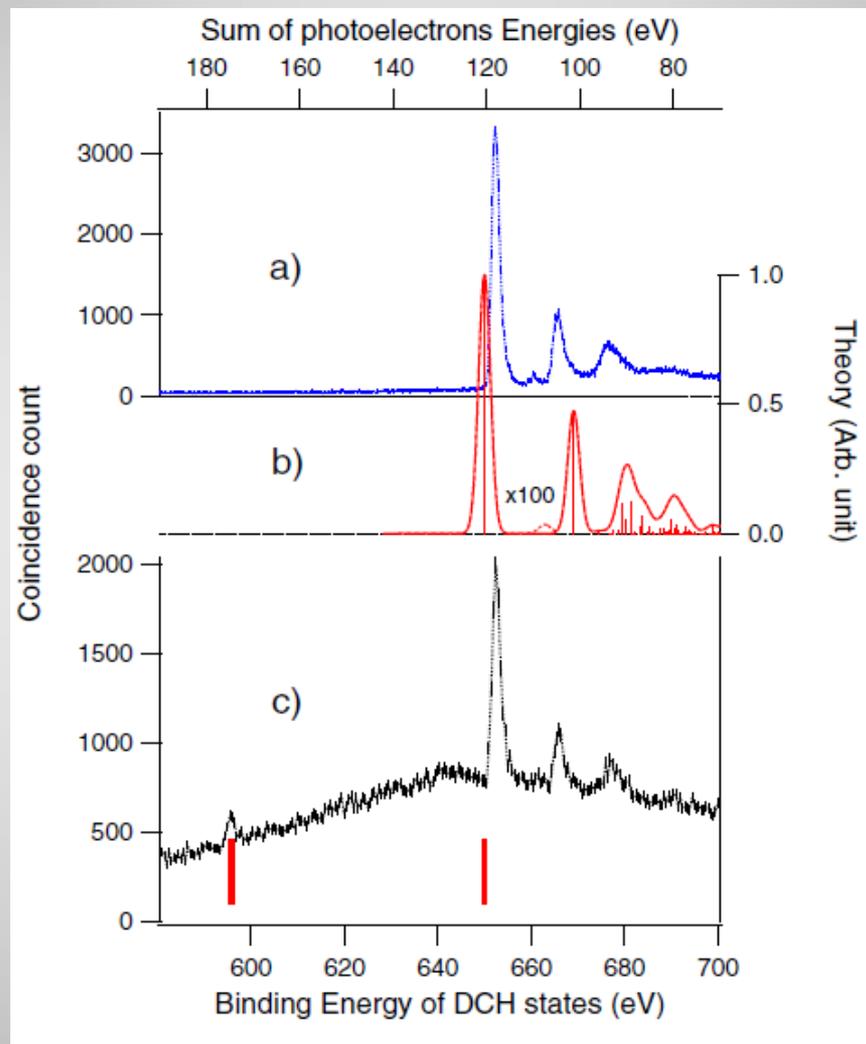


Pulsed light source: He-lamp, SR, fs-laser system, FEL
J.H.D. Eland *et al.*, Phys. Rev. Lett. **90**, 053003 (2003)



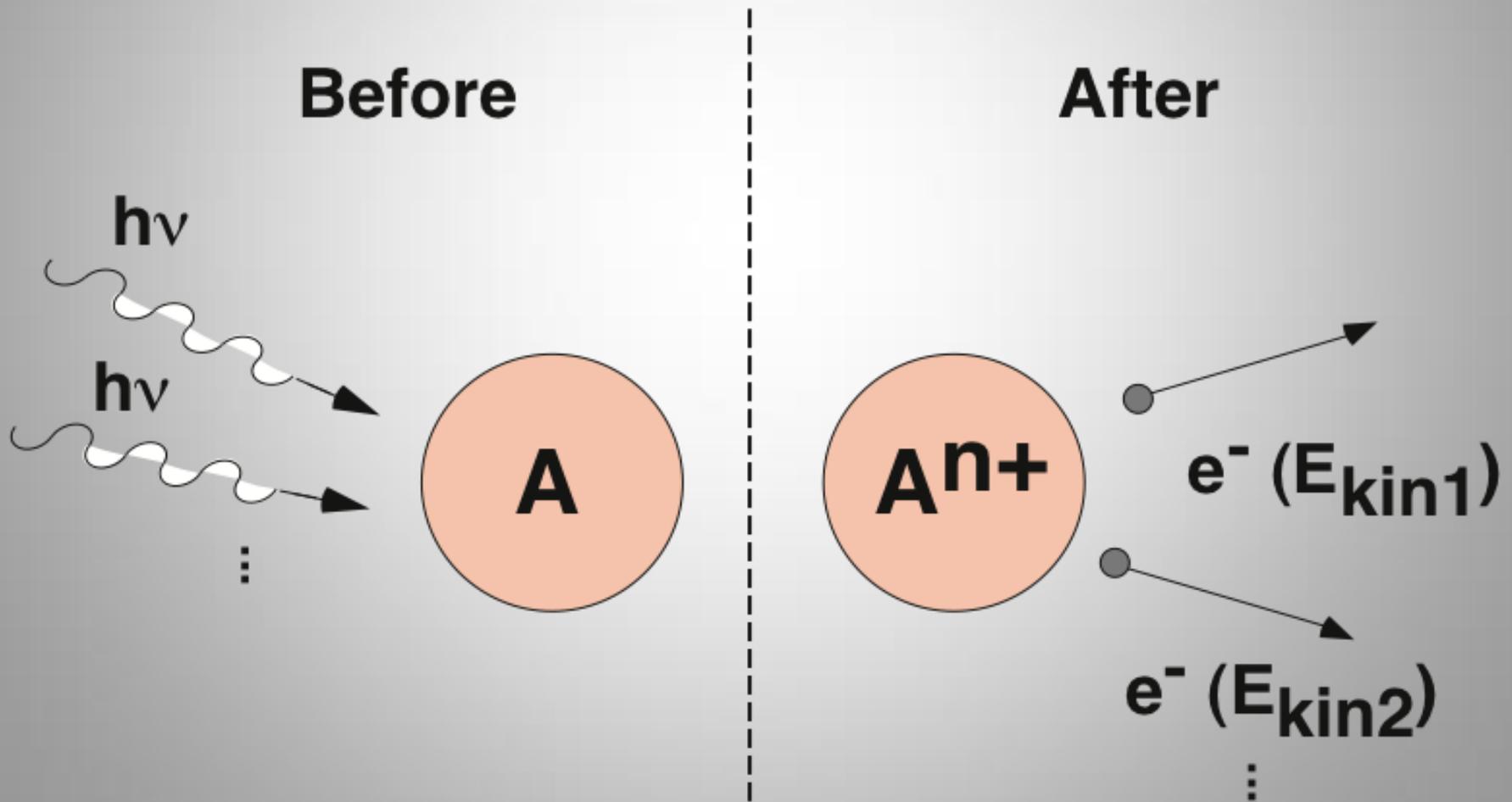
P.Lablanquie, F.Penent, J.Palaudoux, L.Andric, P.Selles, S.Carniato, K.Bučar, M.Žitnik, M. Huttula, J.H.D.Eland, E.Shigemasa, K.Soejima, Y.Hikosaka, I.H.Suzuki, M.Nakano and K.Ito, PRL 106, 063003 (2011)

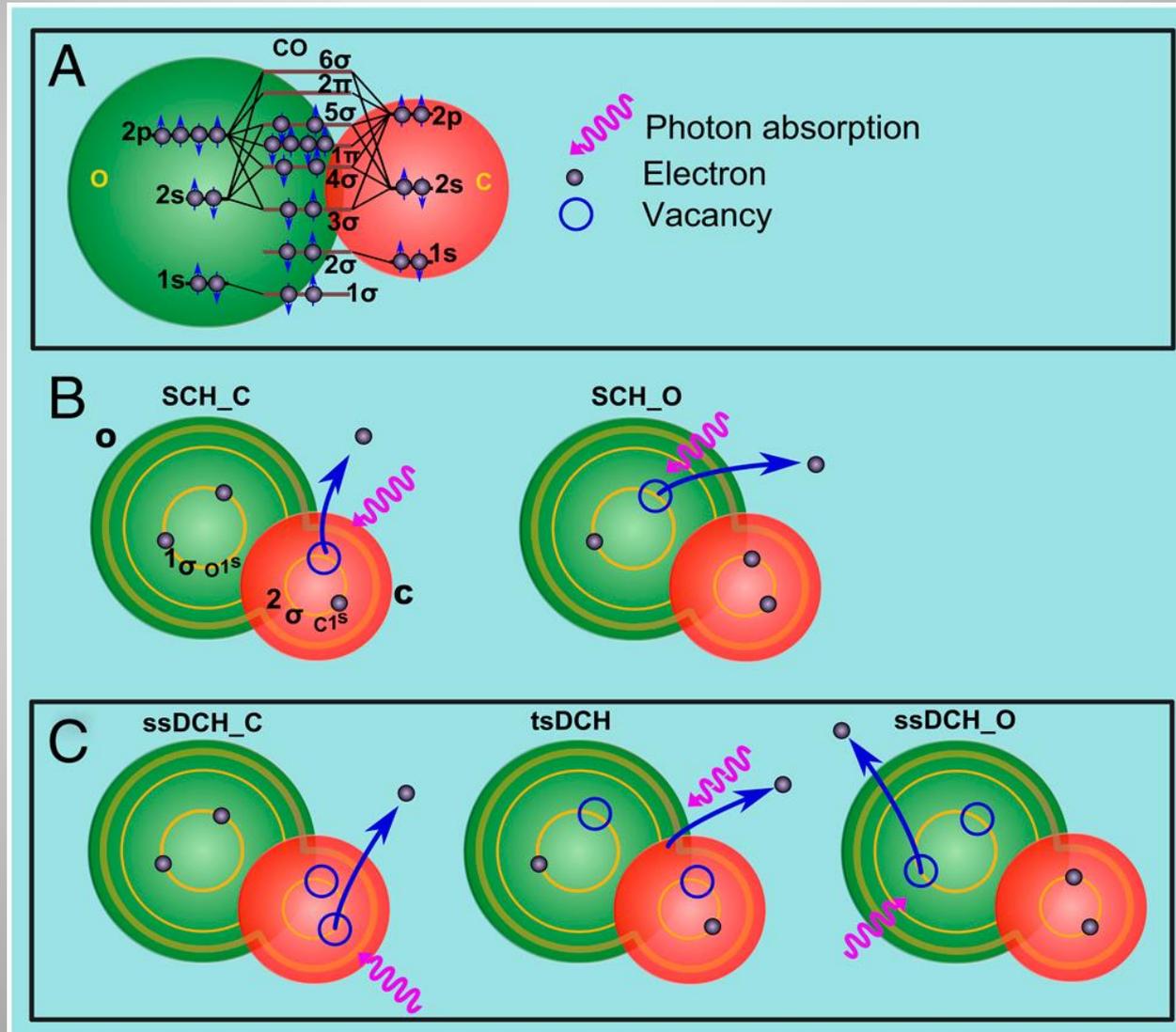






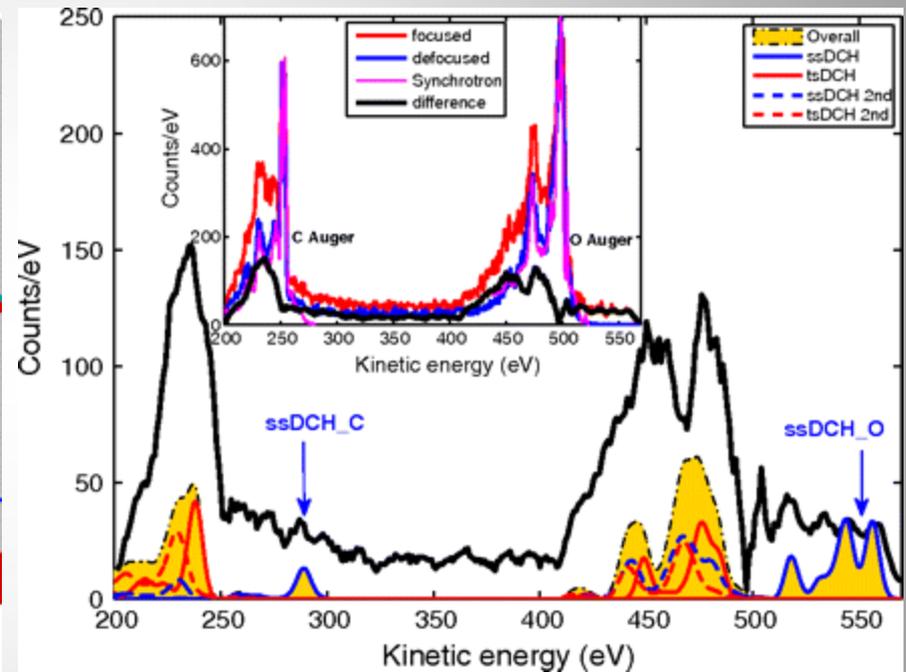
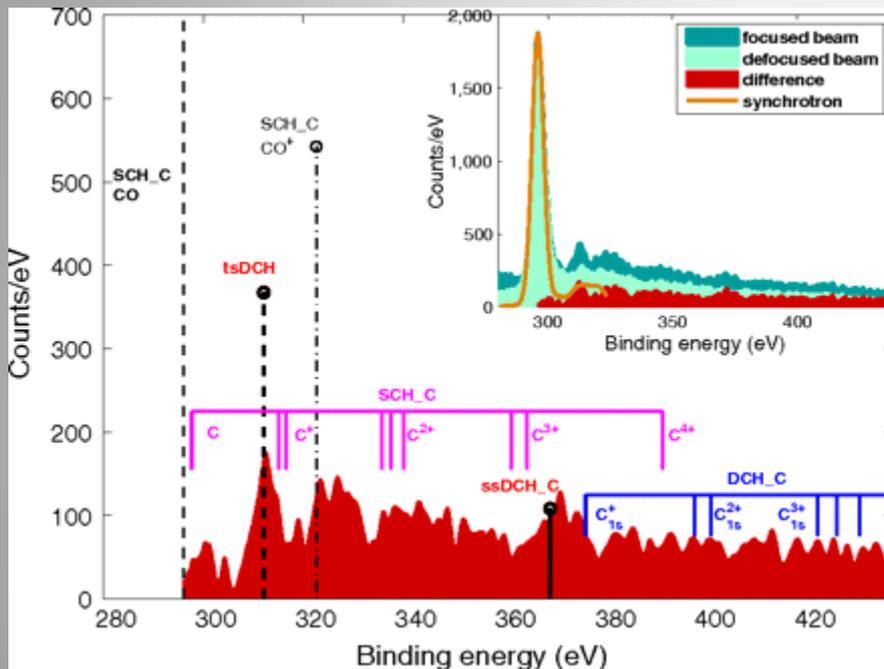
Few Photon – Multiple Ionisation







DCH measurements: CO

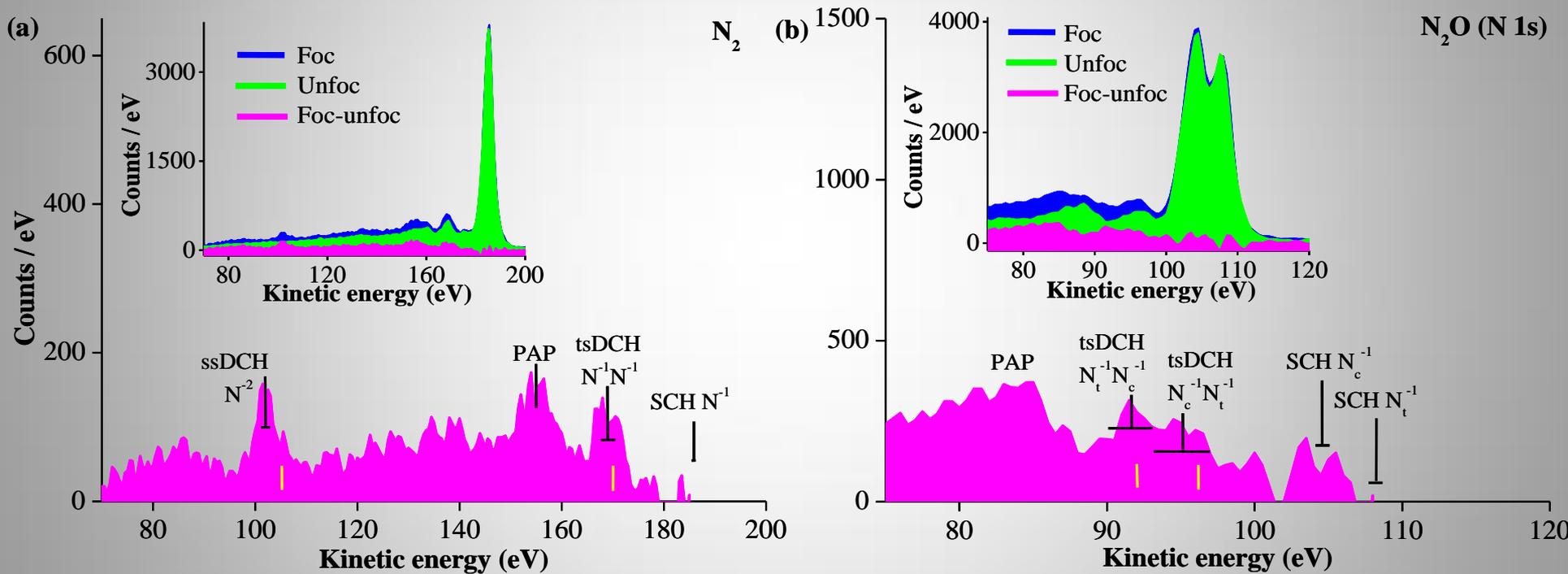


N. Berrah,M.N.Piancastelli *et al.*, PNAS
108, 16912 (2011)

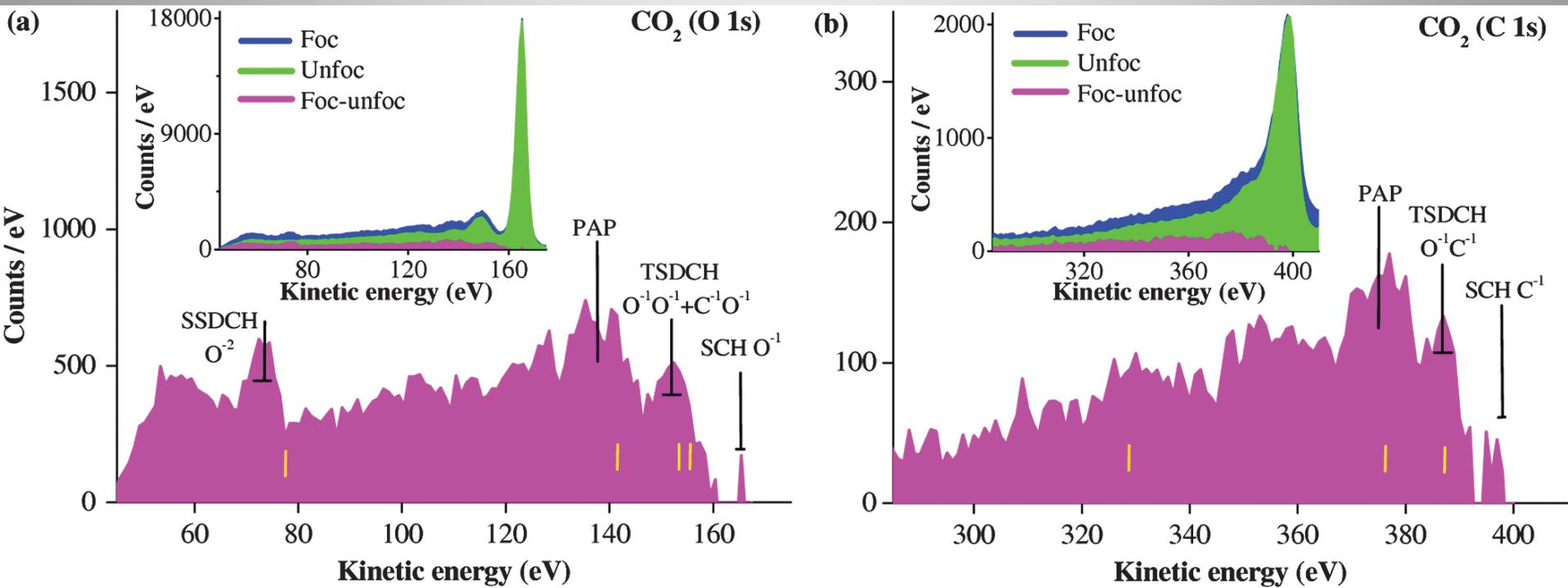
Challenge remains:
TS-DCH → CVV/VVV Auger
and Auger from other channels overlap



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• P. Salén, P. van der Meulen, H.T. Schmidt, R.D. Thomas, M. Larsson, R. Feifel, M.N. Piancastelli, L.Fang, B. Murphy, T. Osipov, N. Berrah, E. Kukk, K. Ueda, J.D. Bozek, C. Bostedt, S. Wada, R. Richter, V. Feyer and K.C. Prince, PRL 108, 153003 (2012)





High performances of the GALAXIES beam line:

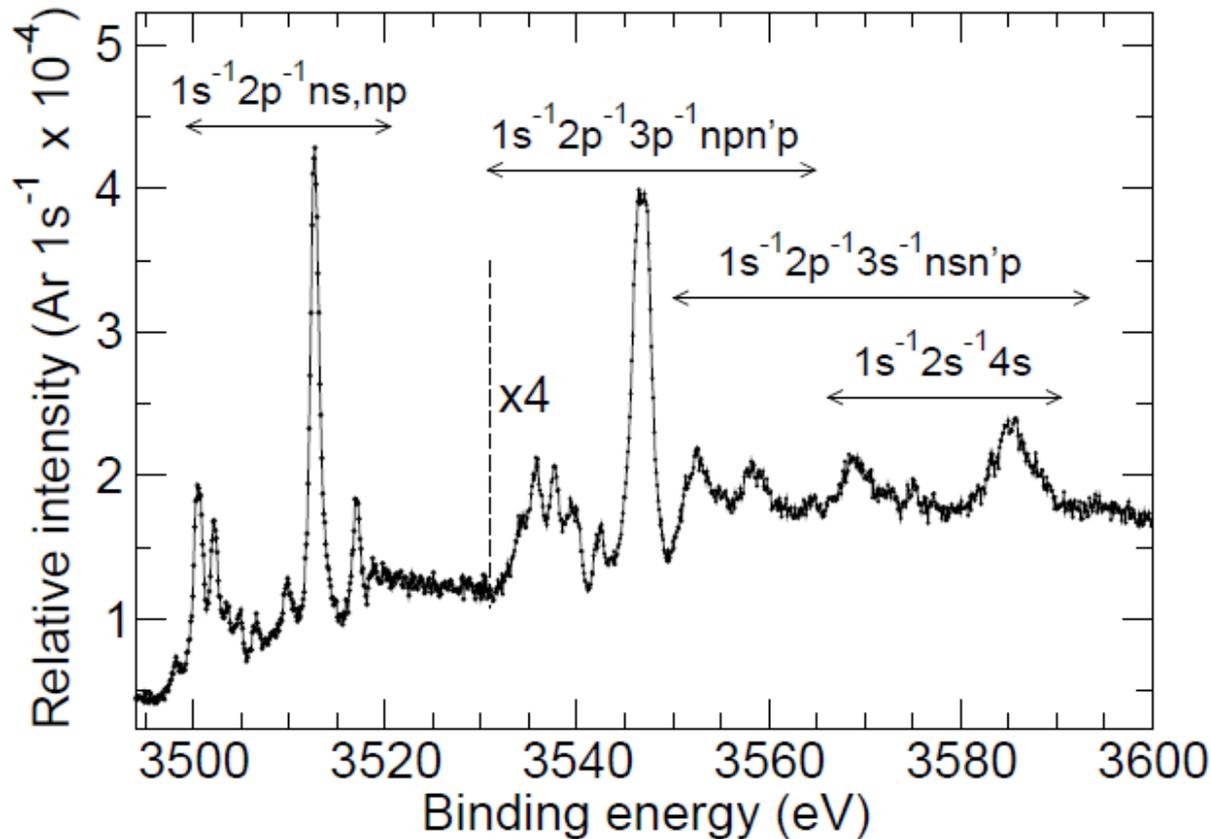
high flux and high resolution

Single-channel measurements

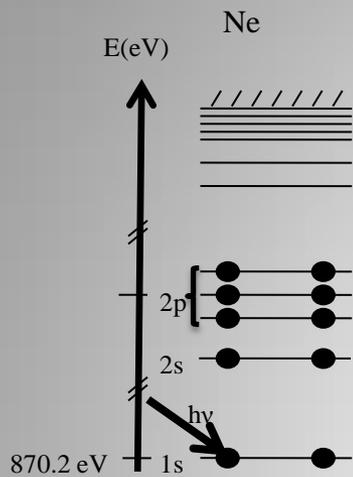
Immediate identification of states of the type:

$$(a) \text{ Ar } 1s^{-1}2p^{-1}np(^2S_{1/2})\epsilon p$$

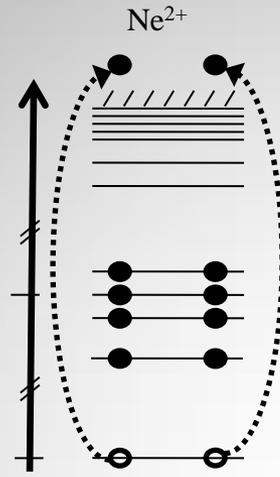
$$(b) \text{ Ar } 1s^{-1}2p^{-1}ns(^2P_{1/2,3/2})\epsilon s, \epsilon d.$$



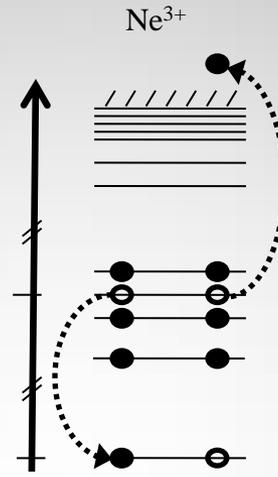
Double core hole (DCH) in Neon



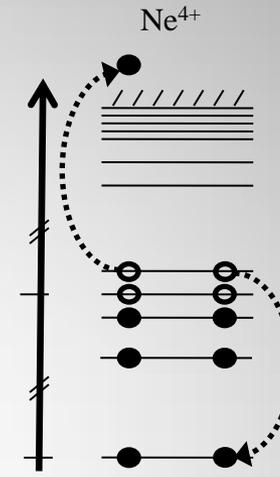
Ground state :
 $1s^2 2s^2 2p^6$



Intermediate state :
 $1s^0 2s^2 2p^6$

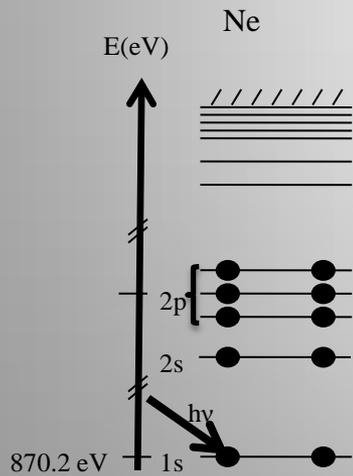


First Auger decay :
 $1s^1 2s^2 2p^4$

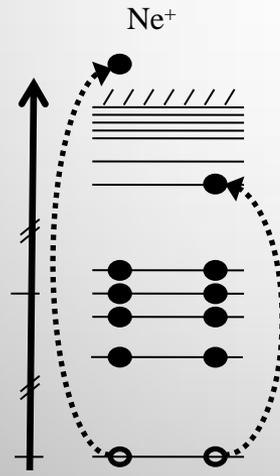


Second Auger decay :
 $1s^2 2s^2 2p^2$

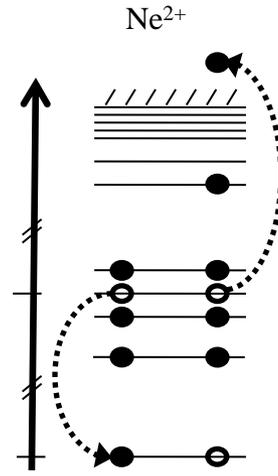
- K^{-2} in Neon.
- Double Ionization Potential = 1863 eV.
- The two photoelectrons share the energy
- First Auger electron = Hypersatellite.
- Second Auger electron = satellite.



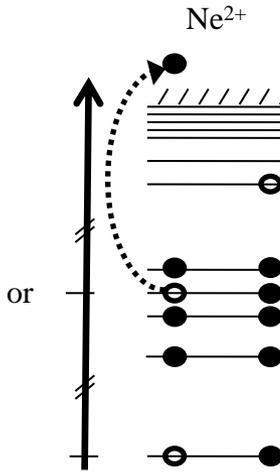
Ground state :
 $1s^2 2s^2 2p^6$



Intermediate state :
 $1s^0 2s^2 2p^6 n l^1$



Spectator Auger decay :
 $1s^1 2s^2 2p^4 n l^1$

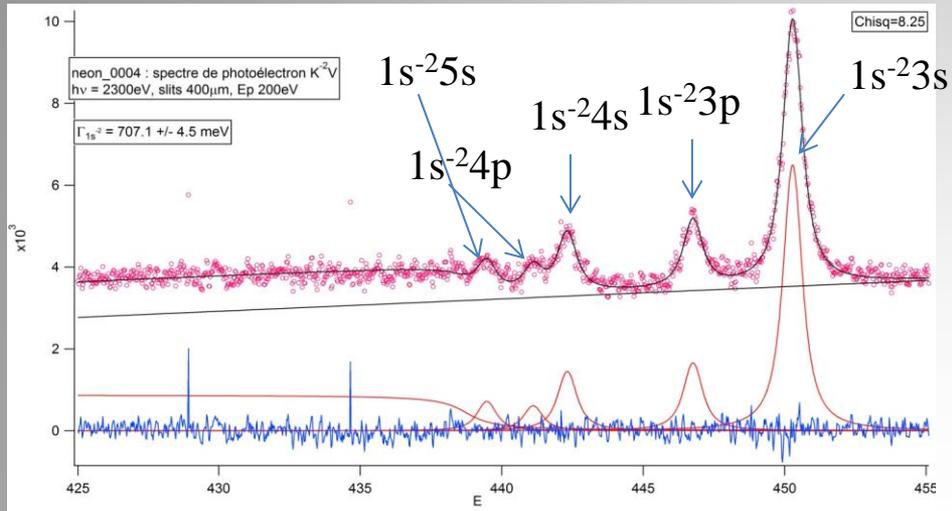


Participator Auger decay :
 $1s^2 2s^2 2p^5$

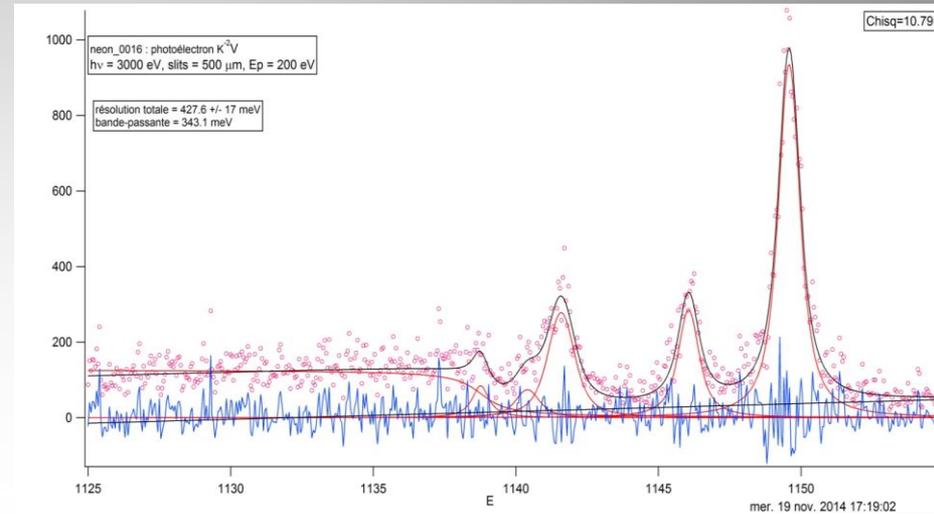
- K^{-2V} in Neon.
- Photoelectrons well separated and easy to identify.
- Different type of decays at different energies.

Photon energy dependence of DCH

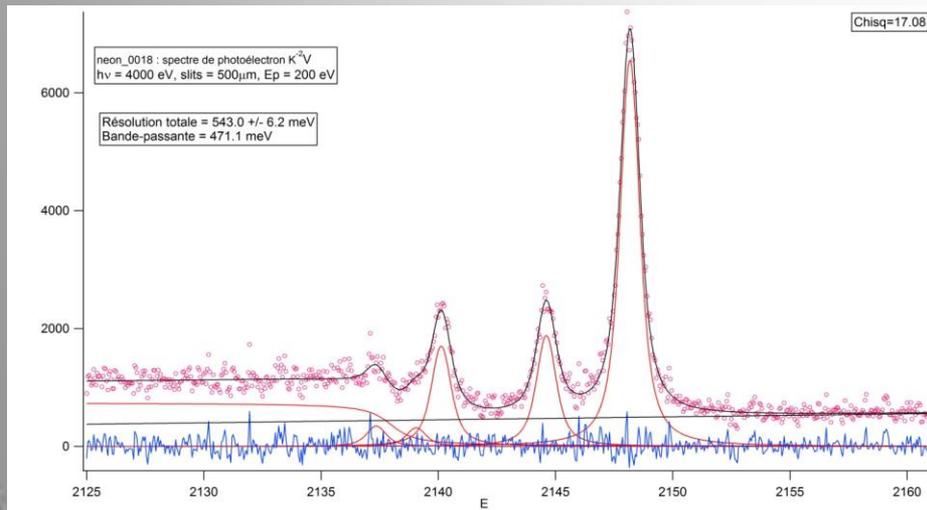
$h\nu = 2300 \text{ eV}$



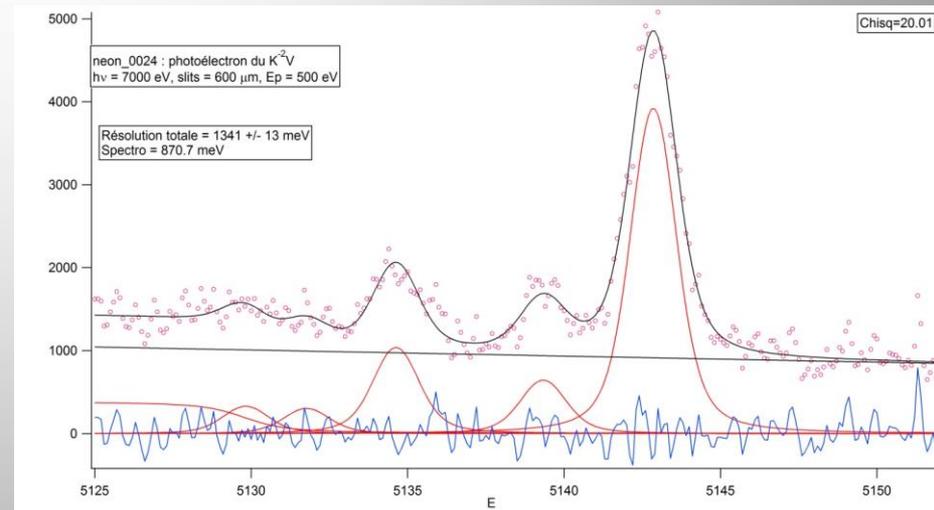
$h\nu = 3000 \text{ eV}$



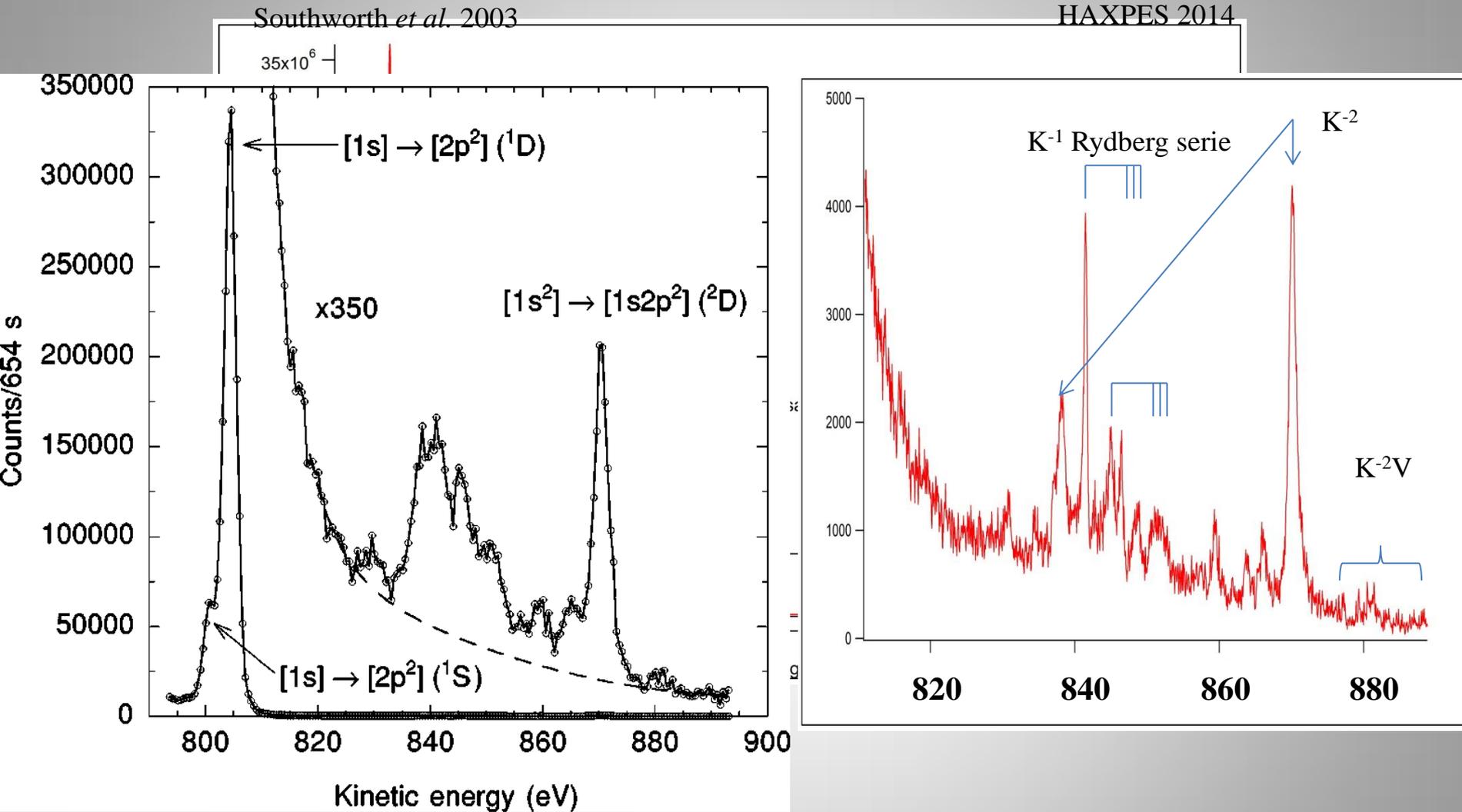
$h\nu = 4000 \text{ eV}$



$h\nu = 7000 \text{ eV}$



The hypersatellite spectra

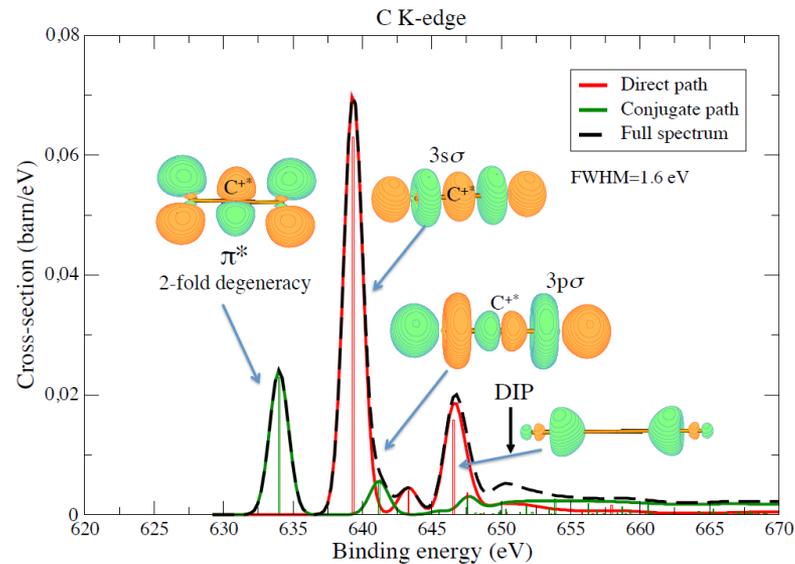
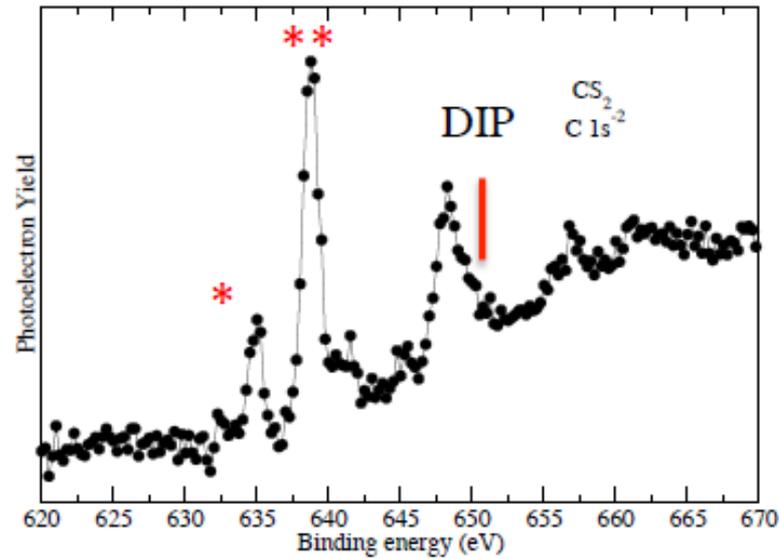


➤ $K^{-1} \approx 1000$ times more intense than K^{-2}

➤ Thanks to much better resolution we were able to identify K^{-1} satellites

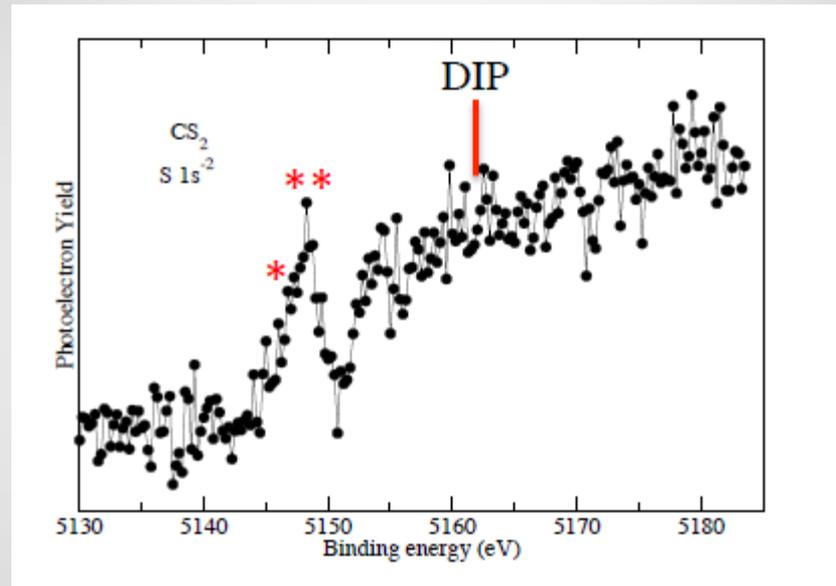
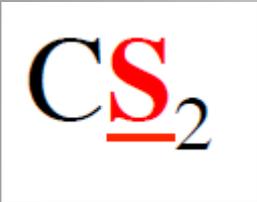
Southworth et al. PRA 67, 062712 (2003)

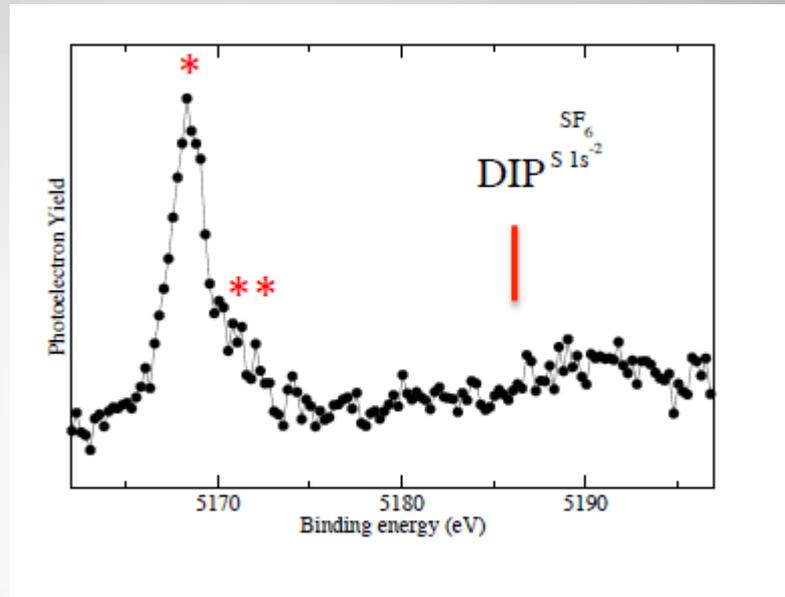
Svensson et al. J. Electron Spectrosc. Relat. Phenom. 47, 327 (1988)



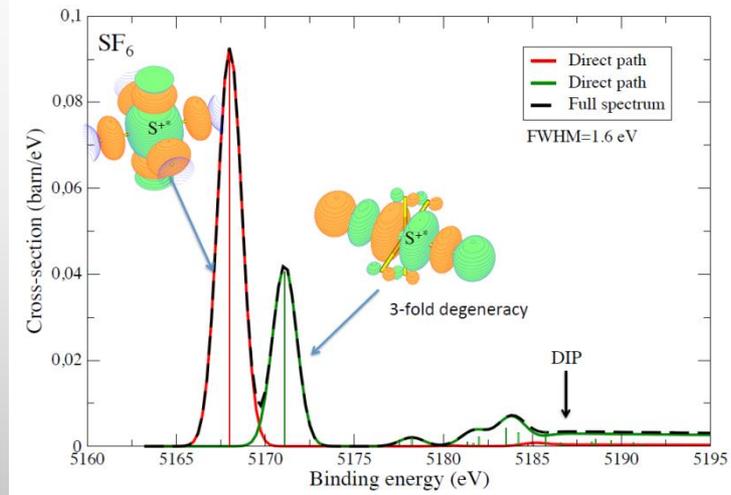


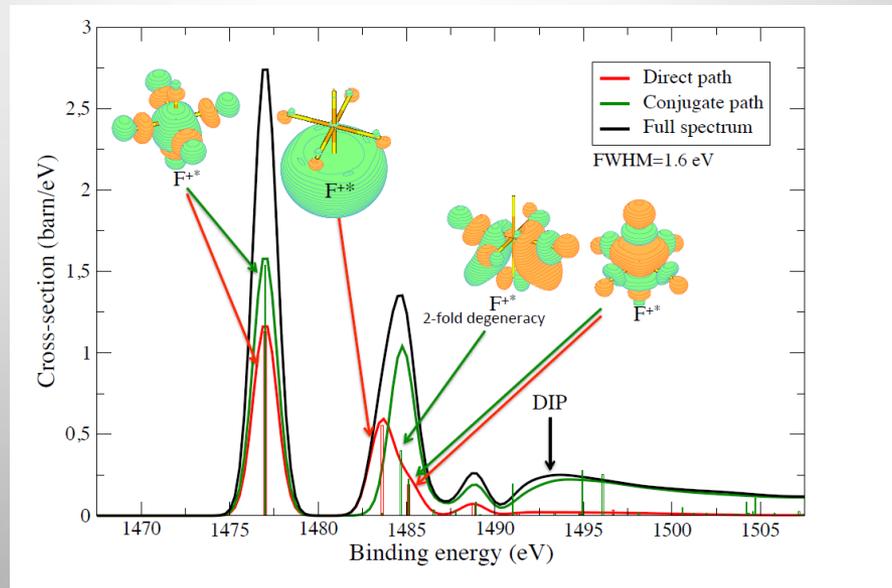
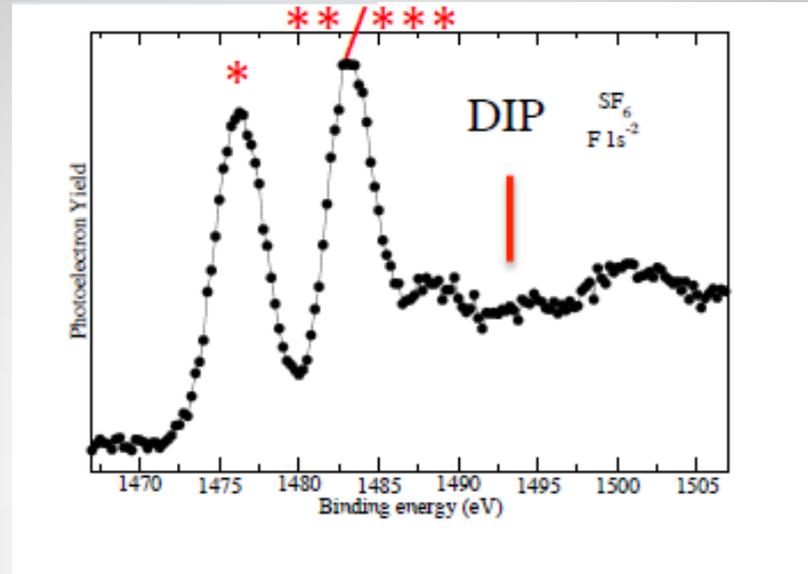
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s-R edge







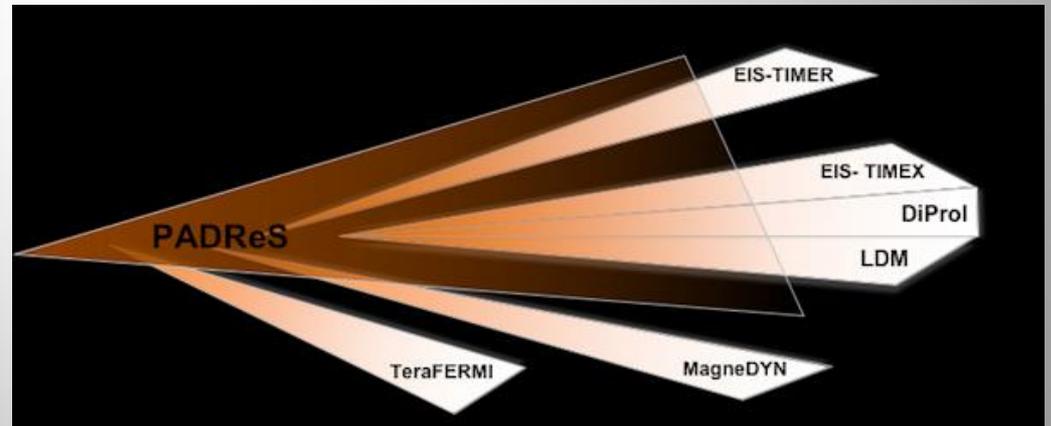
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FERMI

Elettra, Trieste, Italy

Aerial view



Beam transport and beamlines



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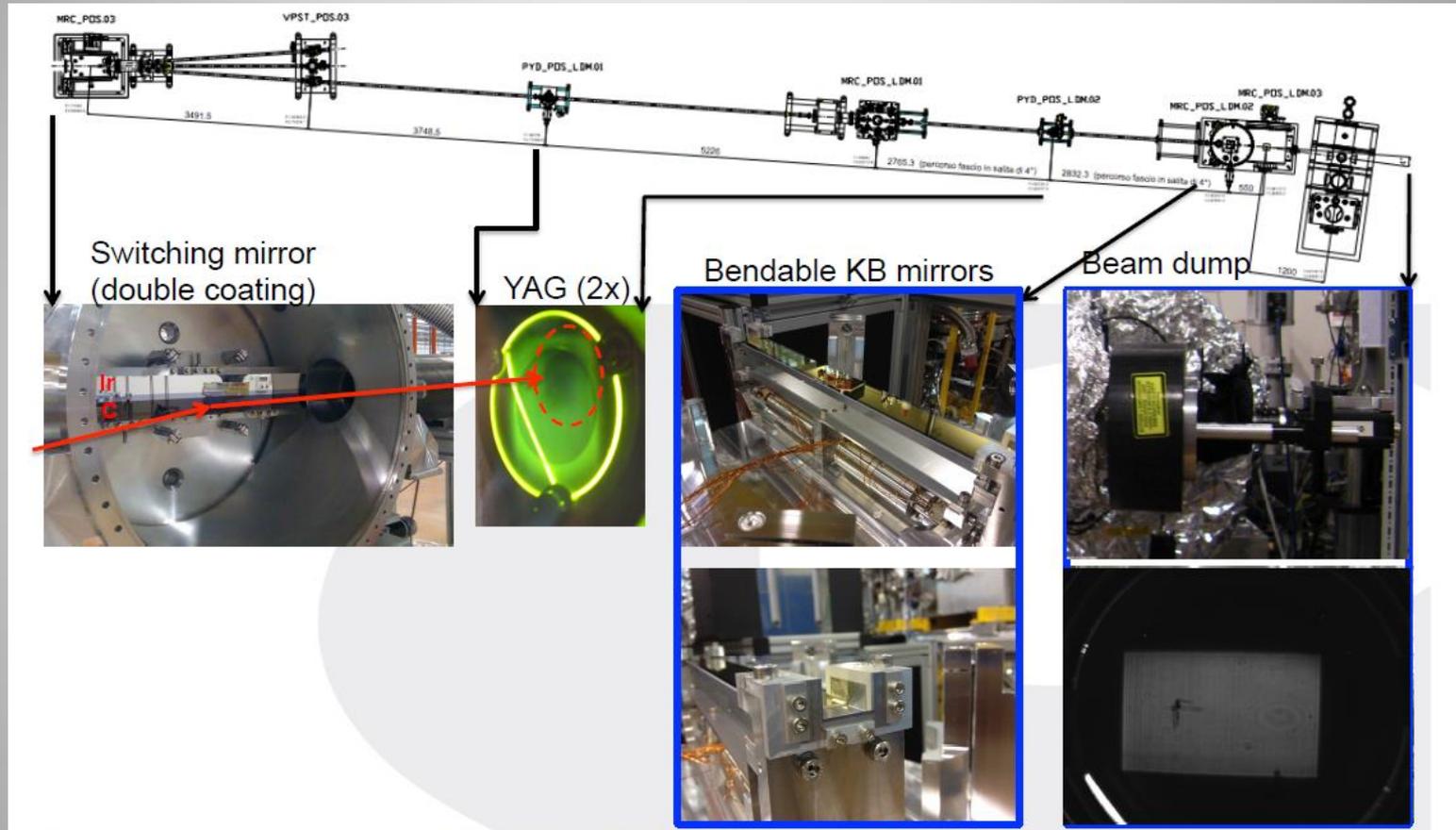
First seeded FEL

	λ seed	261.60	(± 0.5) nm	4.7394	eV	
harmonic stage1↓	1st stage wl (nm)	2nd stage wl (nm)				
4	65.40	21.80	16.35	13.08	10.90	9.34
5	52.32	17.44	13.08	10.46	8.72	7.47
6	43.60	14.53	10.90	8.72	7.27	6.23
7	37.37	12.46	9.34	7.47	6.23	5.34
8	32.70	10.90	8.18	6.54	5.45	4.67
9	29.07	9.69	7.27	5.81	4.84	4.15
10	26.16	8.72	6.54	5.23	4.36	
11	23.78	7.93	5.95	4.76		
12	21.80	7.27	5.45	4.36		
13	20.12	6.71	5.03	4.02		
harmonic stage 2→		3	4	5	6	7
harmonic stage1↓	1st stage En. (eV)	2nd stage Energy (eV)				
4	18.96	56.87	75.83	94.79	113.75	132.70
5	23.70	71.09	94.79	118.49	142.18	165.88
6	28.44	85.31	113.75	142.18	170.62	199.06
7	33.18	99.53	132.70	165.88	199.06	232.23
8	37.92	113.75	151.66	189.58	227.49	265.41
9	42.66	127.97	170.62	213.28	255.93	298.59
10	47.39	142.18	189.58	236.97	284.37	
11	52.13	156.40	208.54	260.67		
12	56.87	170.62	227.49	284.37		
13	61.61	184.84	246.45	308.06		

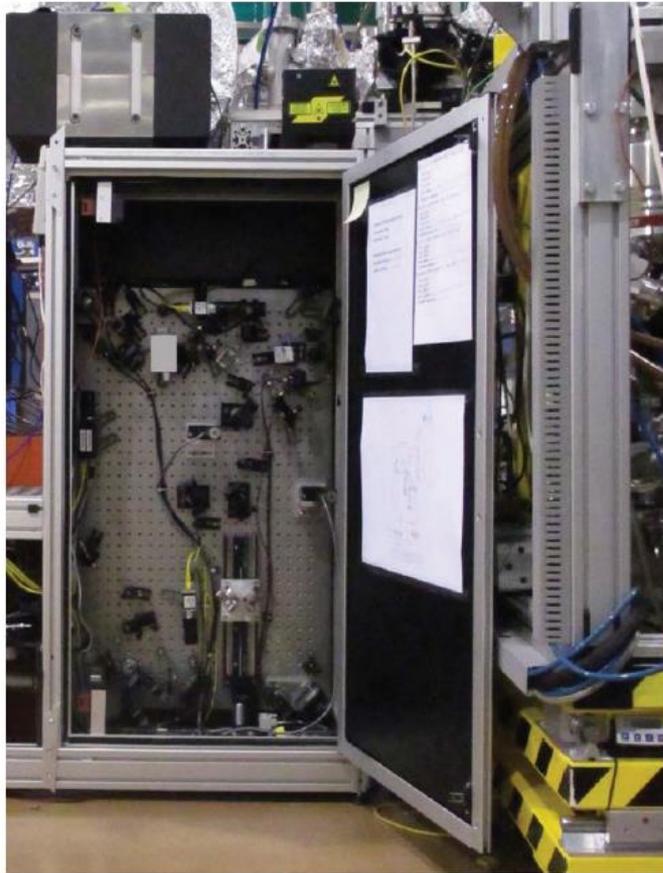
Variable polarization

Negligible photon energy jitter

Negligible time jitter



Low Density Matter (LDM)



Ti:sapphire (same used to form the FEL pulse)

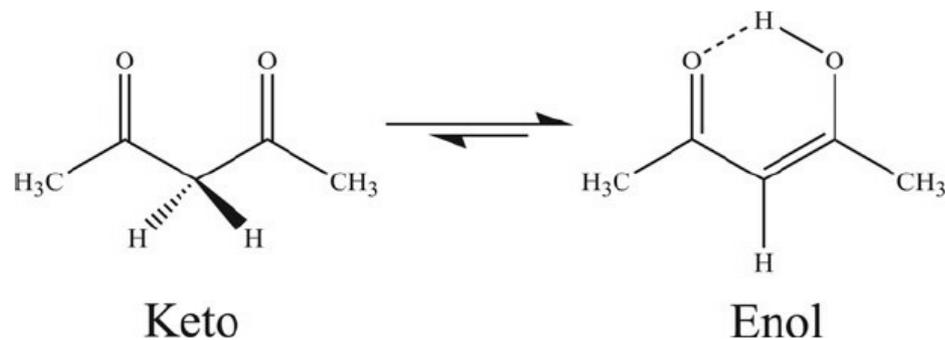
$$\lambda = 780 \text{ nm}$$

Energy/pulse: 1 mJ
Pulse width 140 fs
(autocorrelation)

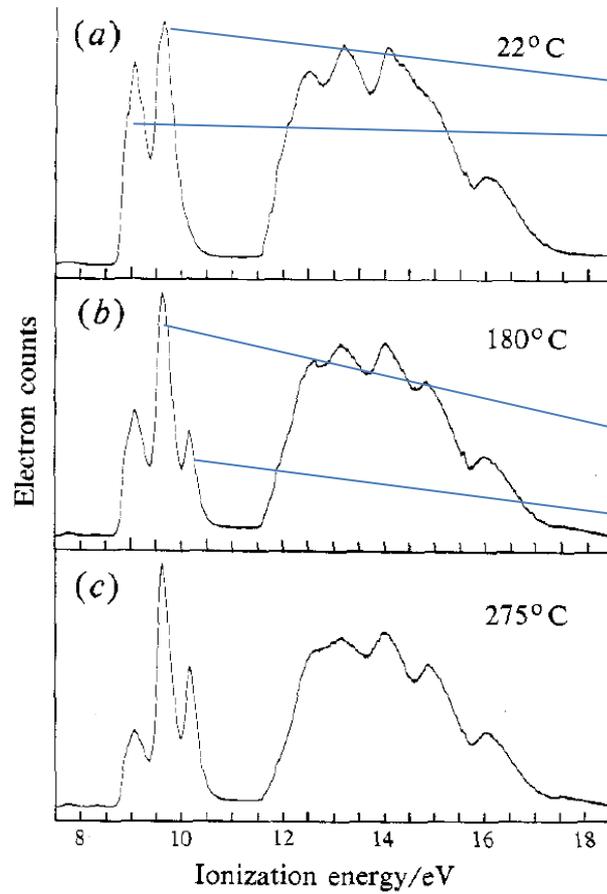
USL-LDM Controls	
Attenuation (energy meter available)	0 – 100 %
Polarization	Full control
Delay resolution	2.6 fs
Beam Position resolution	10 μ rad (upgrade: piezo tip-tilt)



Probing keto-enol tautomerism in acetylacetone



Valence Photoelectron spectra are characteristic of the forms (Hush *et al.* Aust. J.Chem. **40**, (1987) 559)

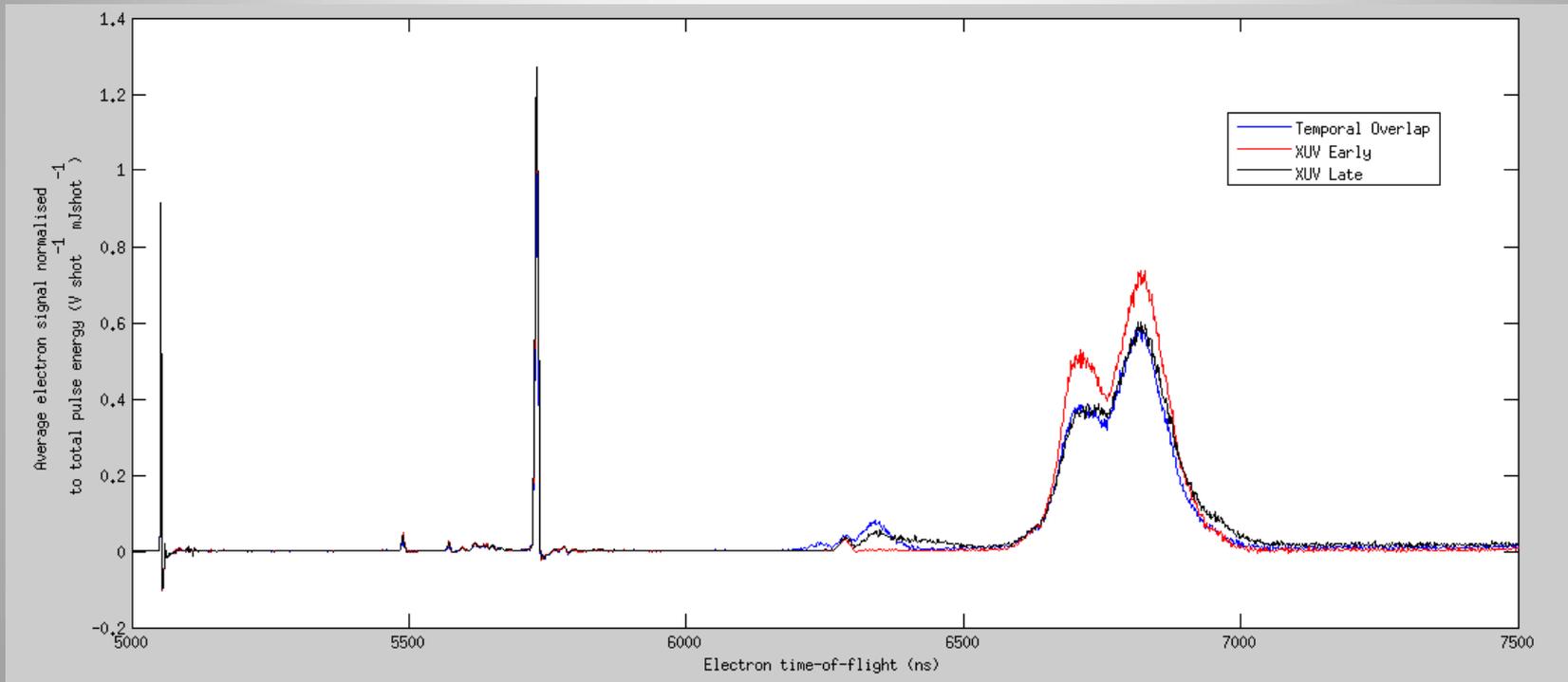


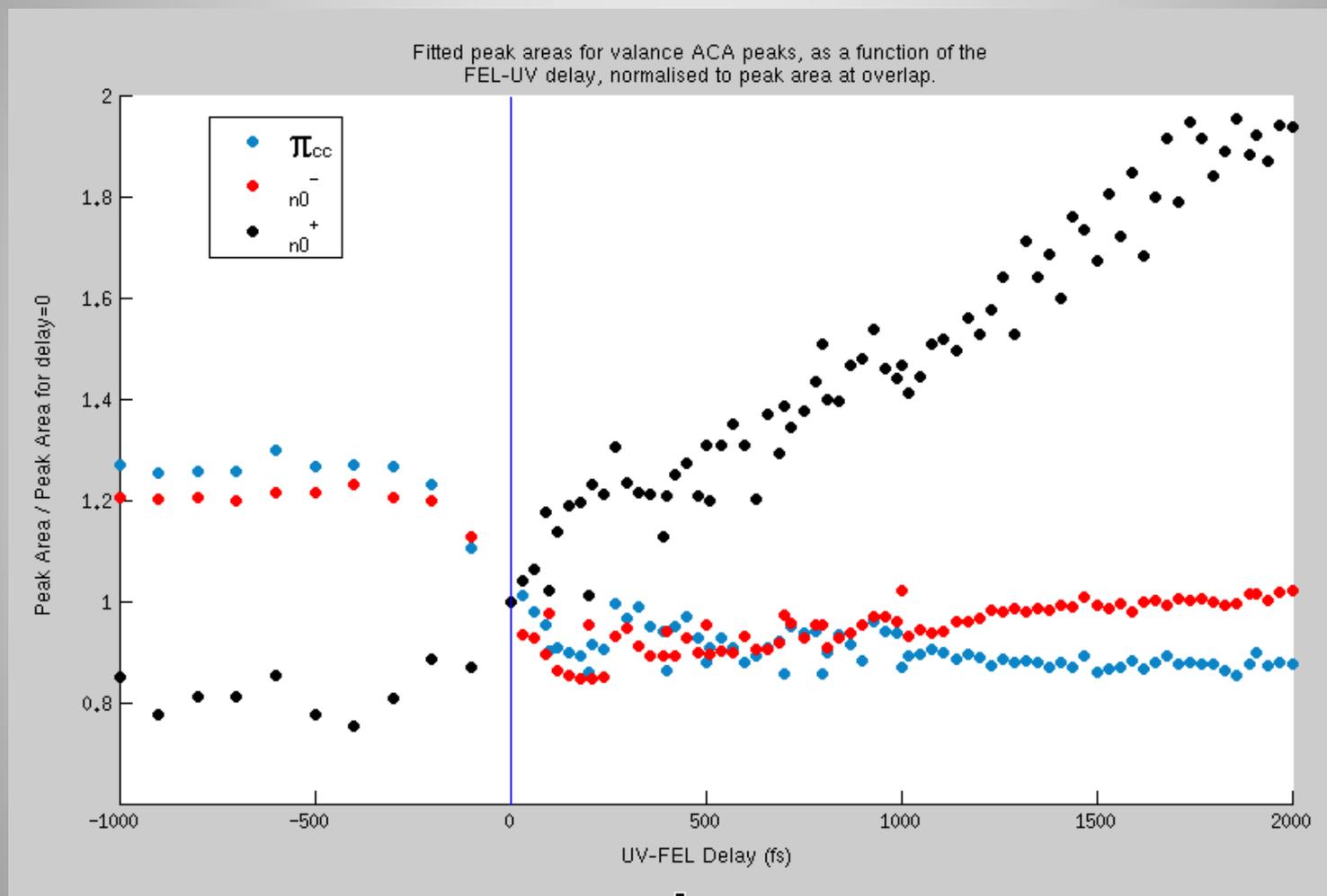
Enol form

Keto form



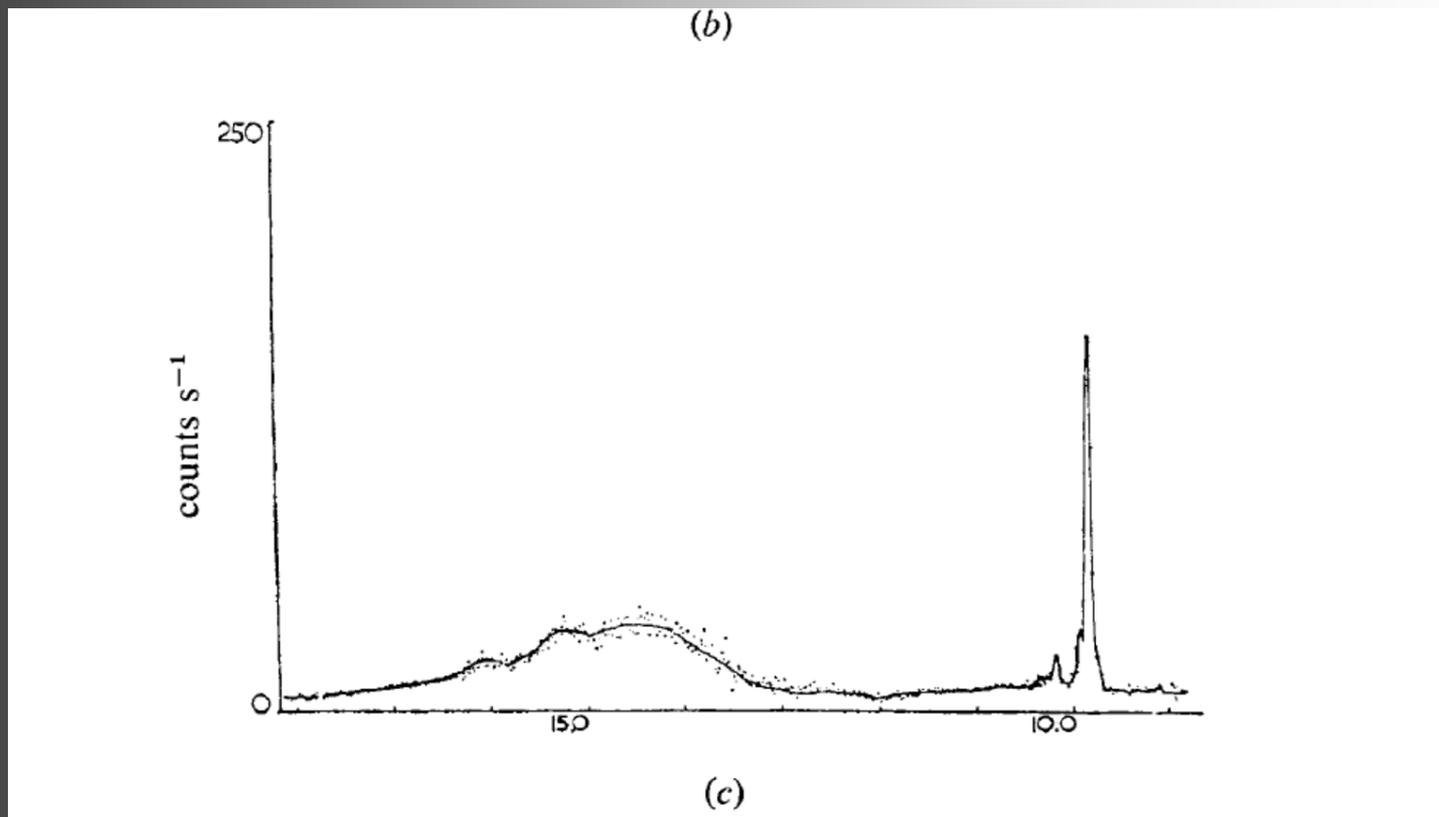
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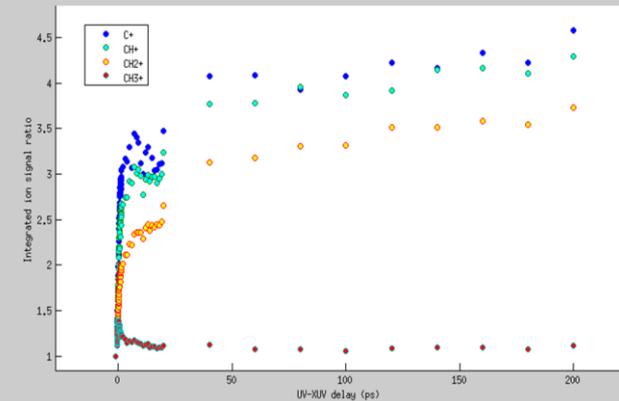
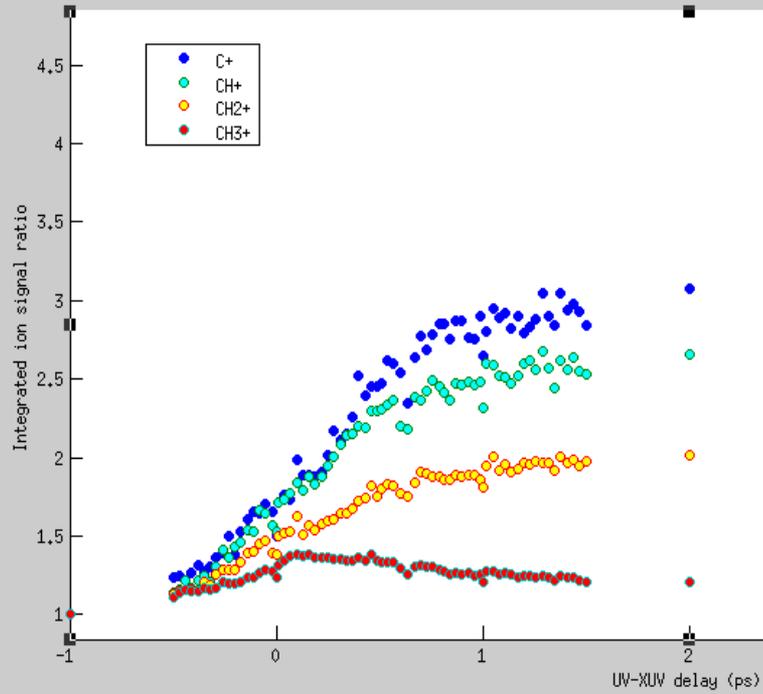
CH₃



J. Dyke, N. Jonathan, E. Lee and A. Morris
J. Chem. Soc., Faraday Trans. 2 72, (1976) 1385

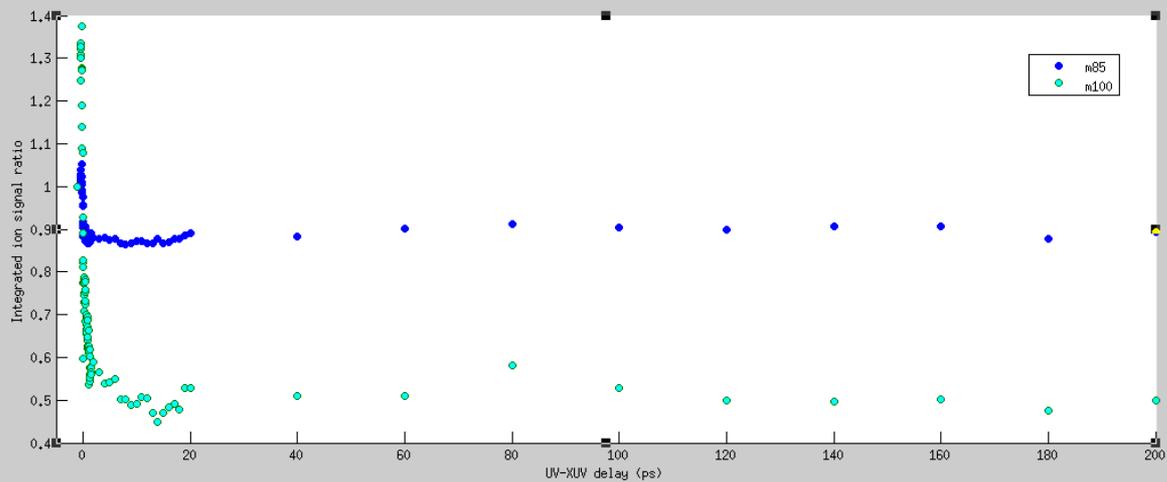
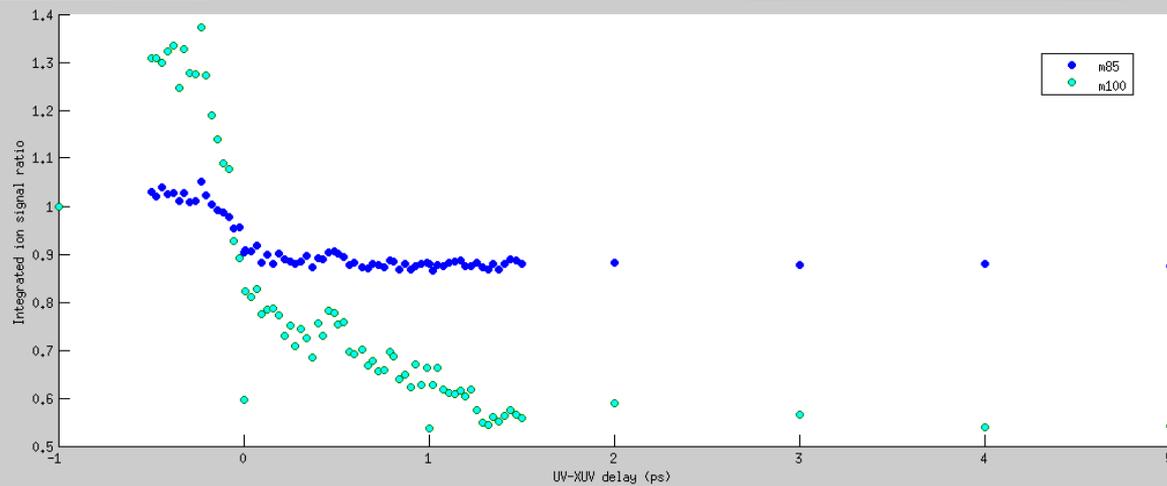


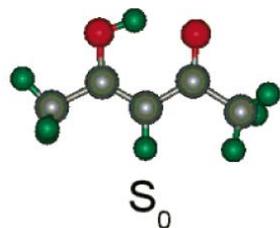
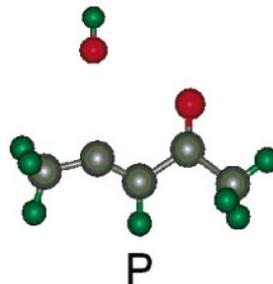
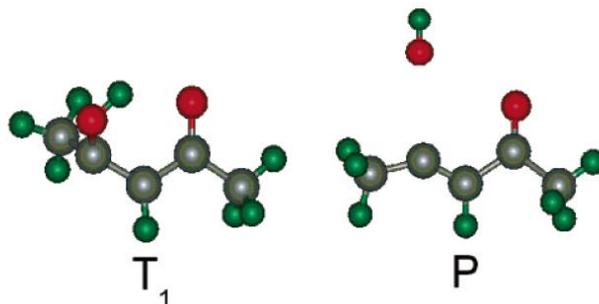
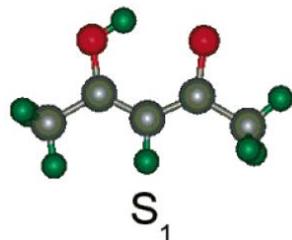
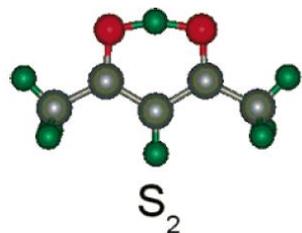
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J. Phys. Chem. A **2004**, *108*, 6650–6655

Ultrafast Electron Diffraction: Structural Dynamics of the Elimination Reaction of Acetylacetone

Shoujun Xu, Sang Tae Park, Jonathan S. Feenstra, Ramesh Srinivasan, and Ahmed H. Zewail*

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